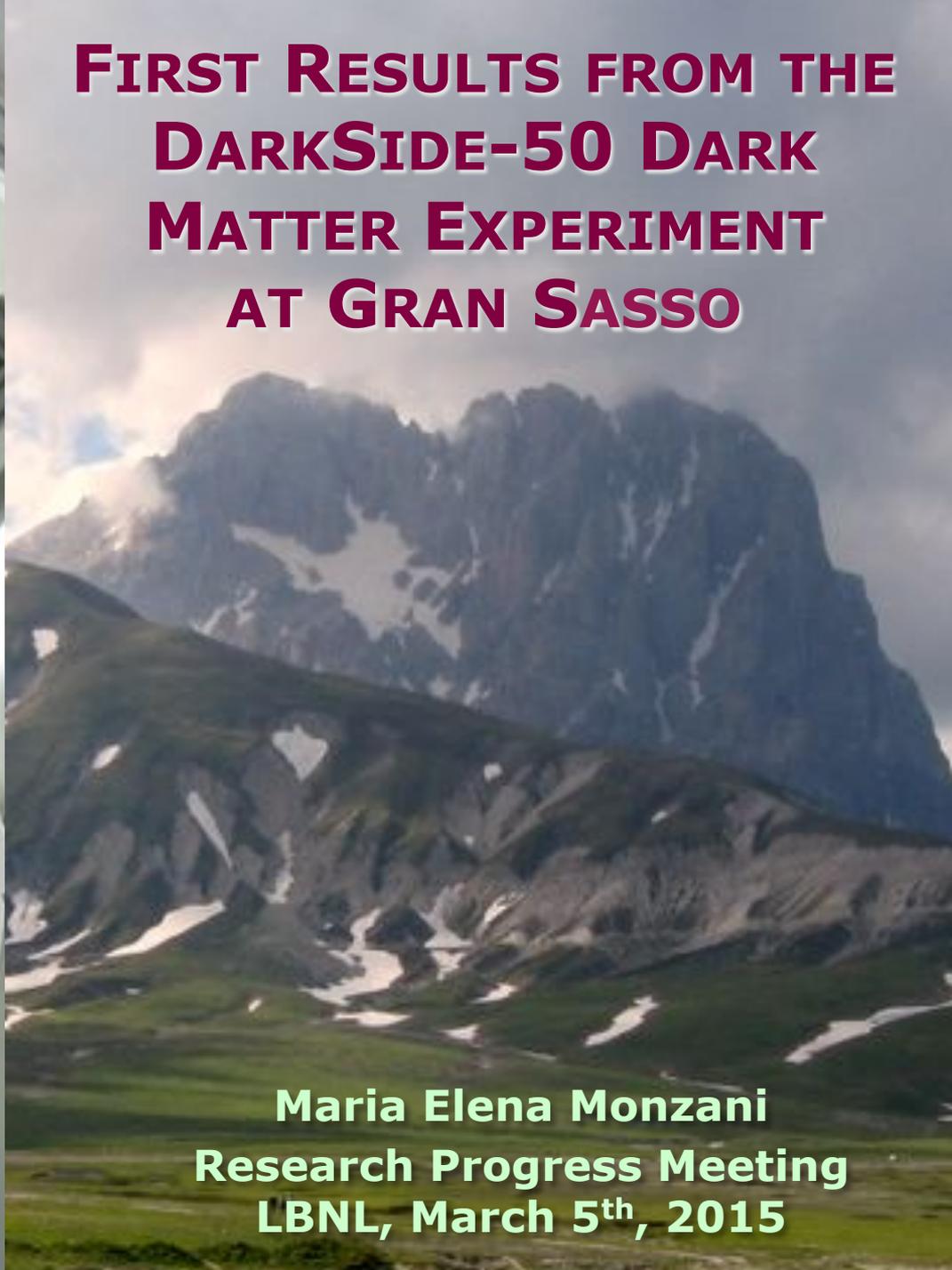


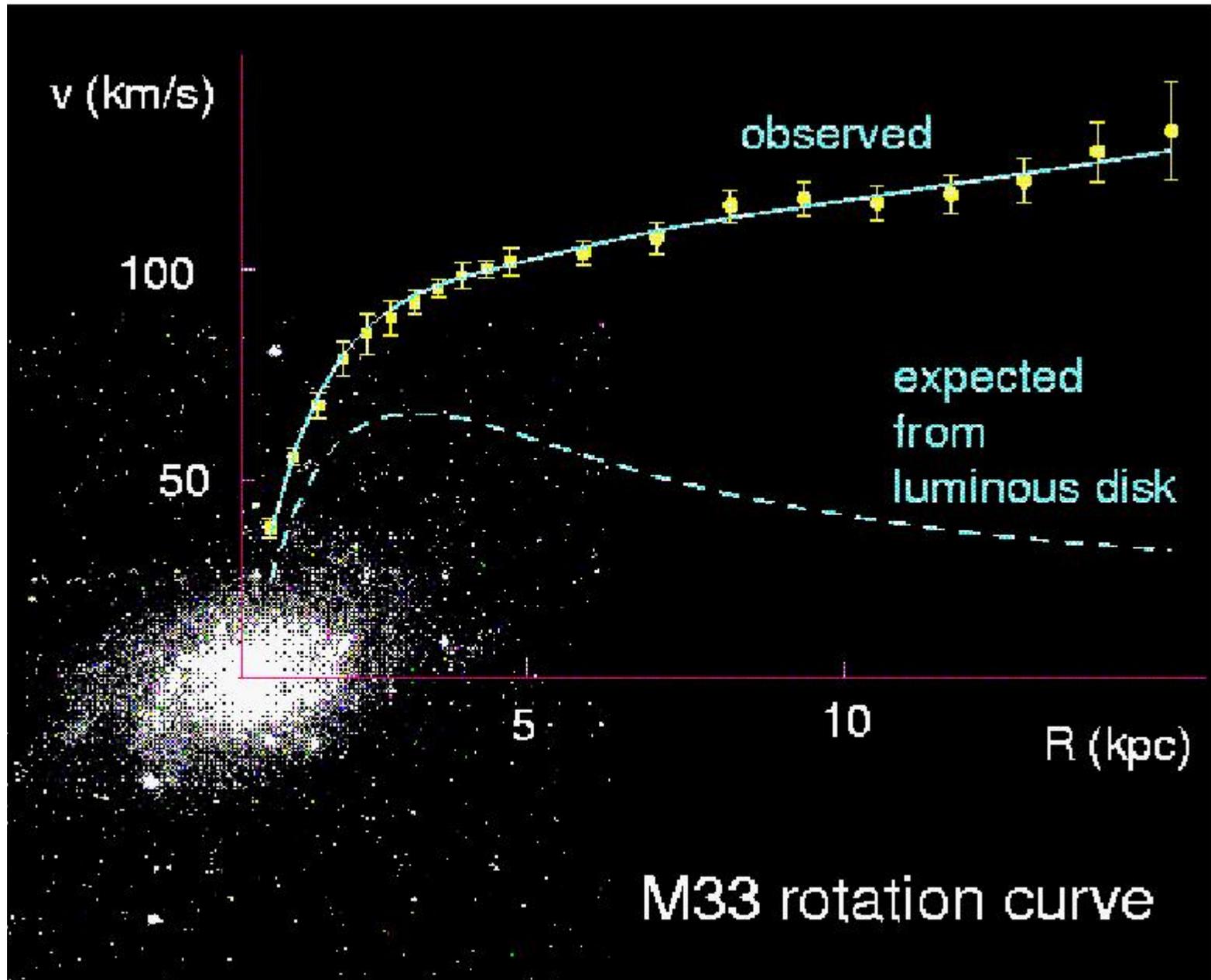


# FIRST RESULTS FROM THE DARKSIDE-50 DARK MATTER EXPERIMENT AT GRAN SASSO

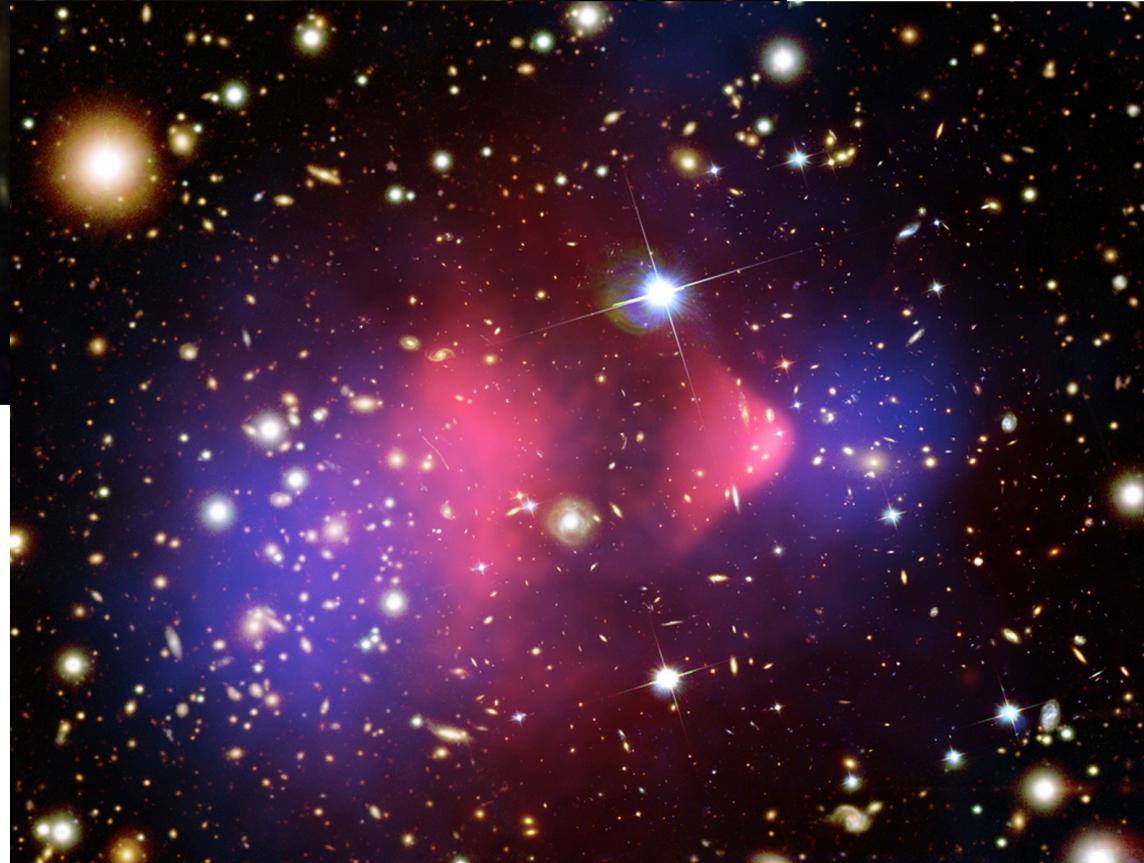


**Maria Elena Monzani**  
**Research Progress Meeting**  
**LBLN, March 5<sup>th</sup>, 2015**

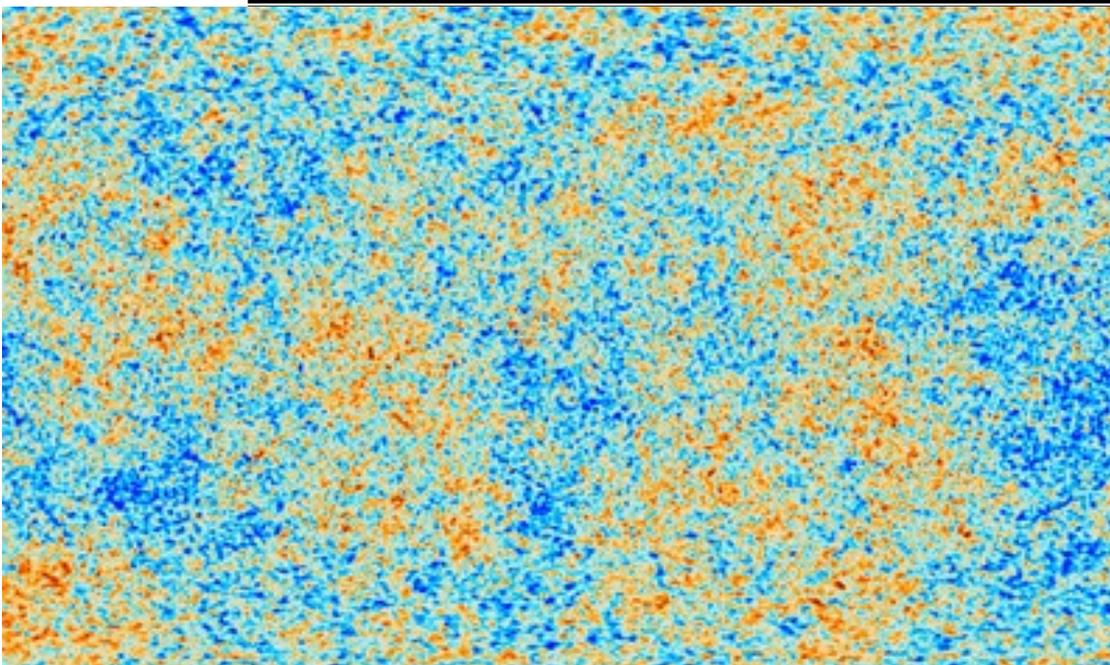
# SOMETHING FISHY WITH SPIRAL GALAXIES



# SOMETHING FISHY WITH GALAXY CLUSTERS



# SOMETHING FISHY WITH THE UNIVERSE



CMB anisotropies:

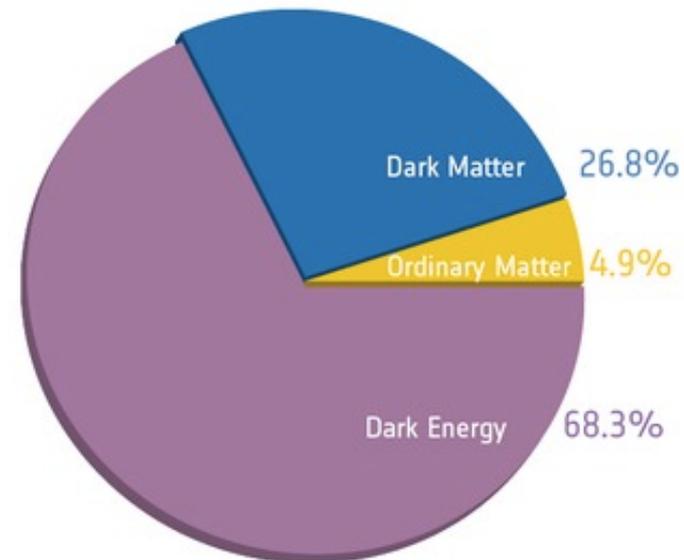
Angular Scale:  $\Omega_{\text{tot}} \approx 1 \Rightarrow \Omega_{\Lambda} + \Omega_{\text{M}} \approx 1$

Odd-vs-even peaks:  $\Omega_{\text{B}} / \Omega_{\text{tot}} \approx 4.9\%$

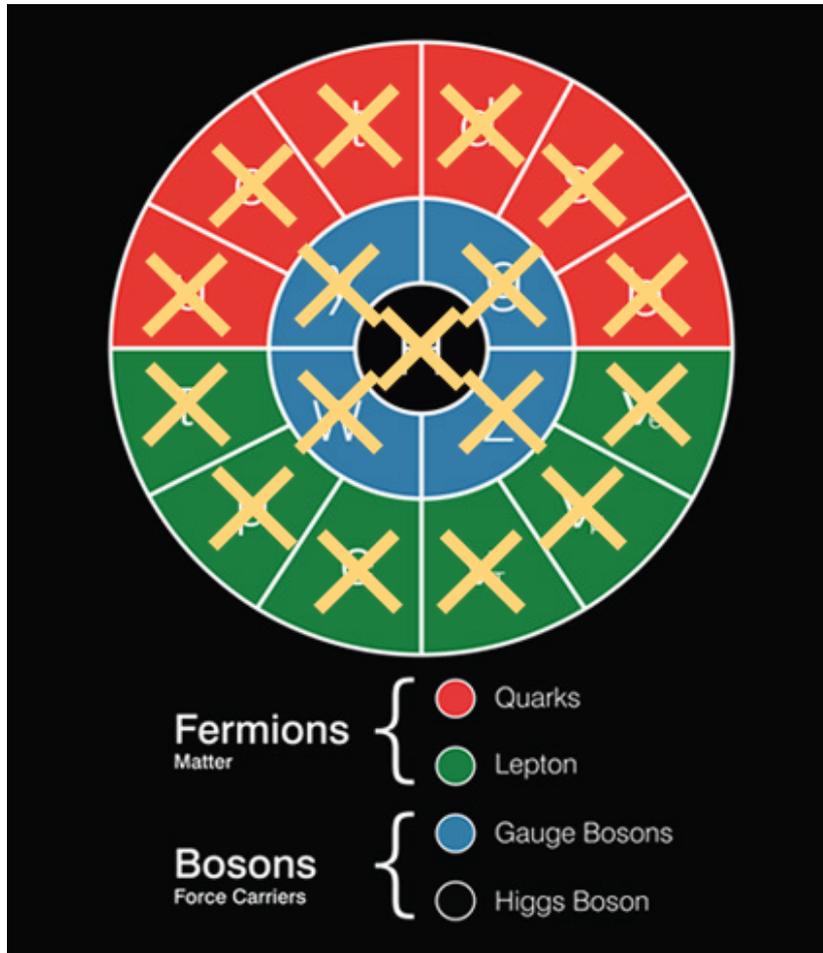
Type Ia SN:

$\Omega_{\Lambda} \approx 1.33 \Omega_{\text{M}} + 0.33 \Rightarrow \Omega_{\text{M}} / \Omega_{\text{tot}} \approx 32\%$

$\Rightarrow \Omega_{\text{DM}} / \Omega_{\text{tot}} \approx 27\%$



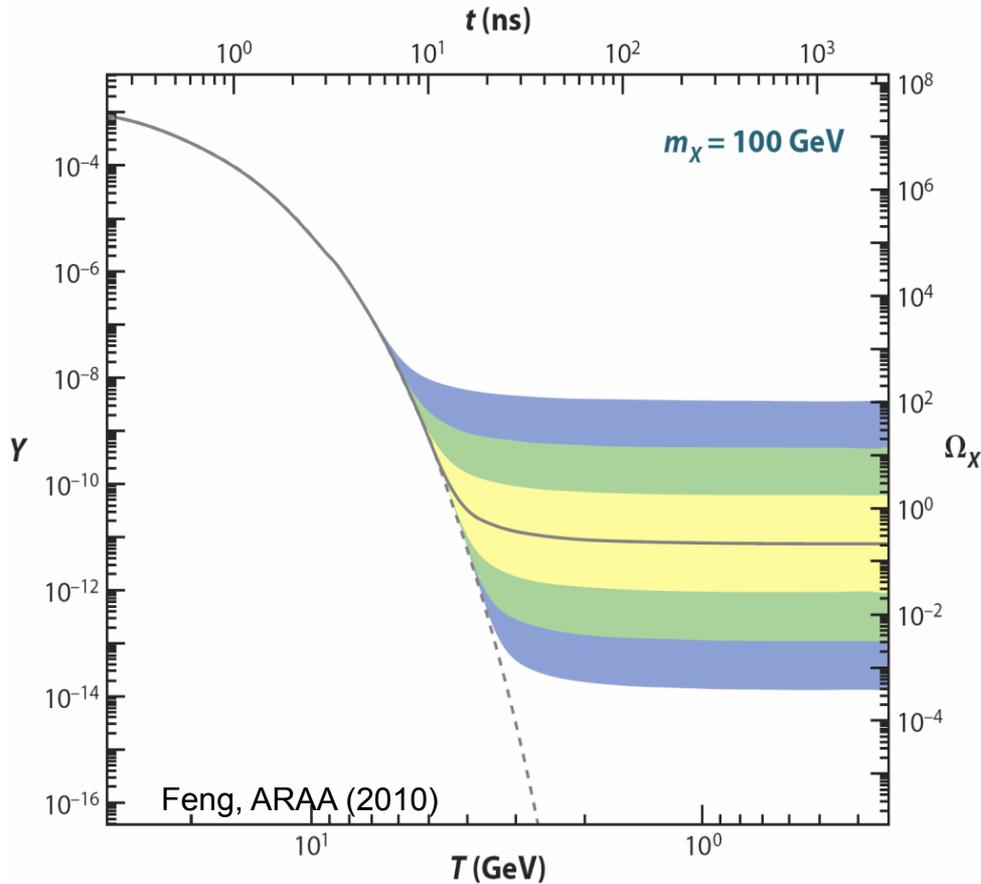
# KNOWN DARK MATTER PROPERTIES



- Gravitationally interacting
- Not short-lived
- Not hot
- Not baryonic

**UNAMBIGUOUS EVIDENCE FOR NEW PARTICLES**

# THE WIMP MIRACLE



If Dark Matter particles are relics from the Big-Bang, the amount of Dark Matter left over from the universe cool-down is determined by its annihilation strength:

$$\Omega_X \propto \frac{1}{\langle \sigma v \rangle} \sim \frac{m_X^2}{g_X^4}$$

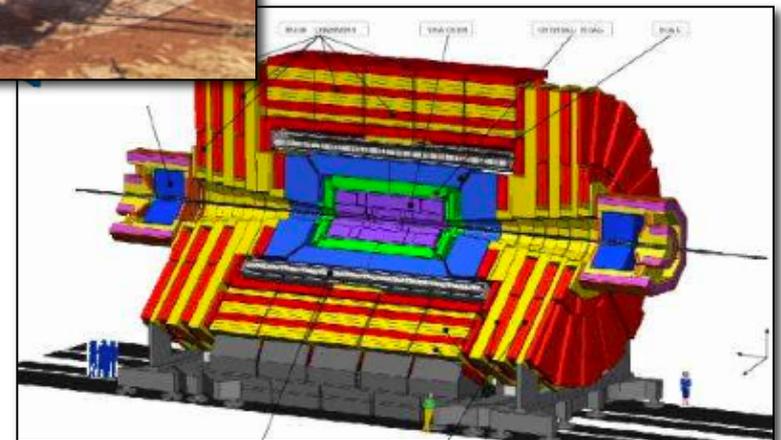
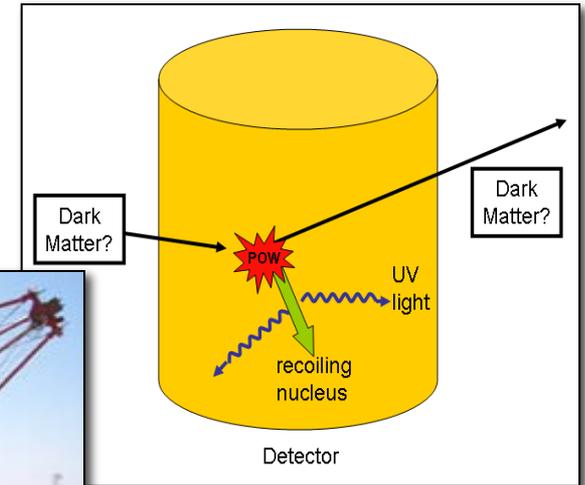
$$\Omega_\chi \sim 0.1 \rightarrow \langle \sigma v \rangle \sim 1 \text{ pb}$$

$$m_\chi \sim 100 \text{ GeV} - 1 \text{ TeV}$$

Remarkable coincidence: particle physics independently predicts particles with the right density to be dark matter

# STRATEGIES FOR DARK MATTER HUNTING

- Direct Detection
- Indirect Detection
- Collider Searches

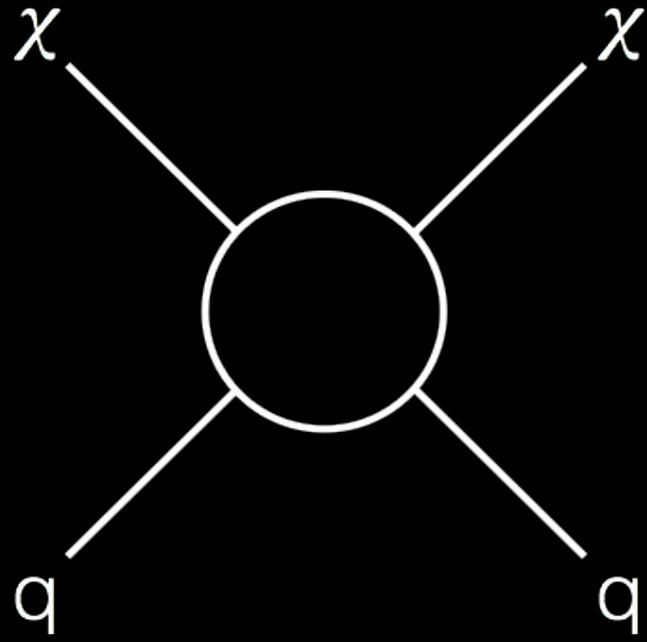


# 3 WAYS TO DETECT WIMPS

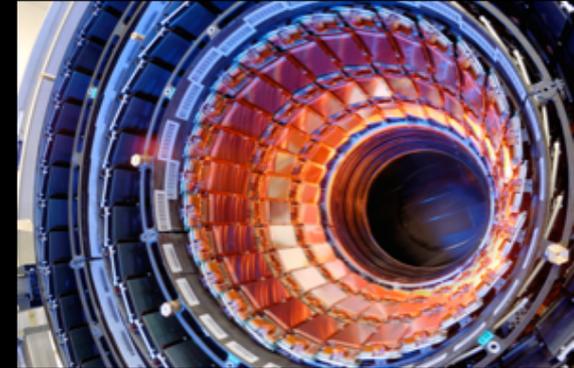
Annihilation



Indirect Detection

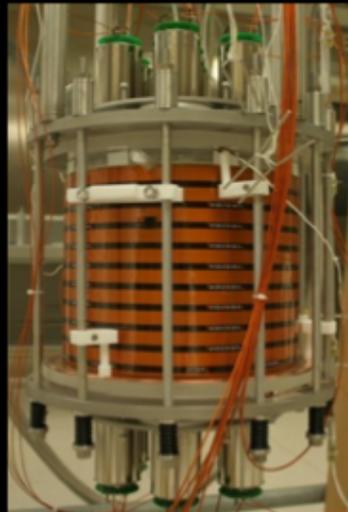


Production



Colliders

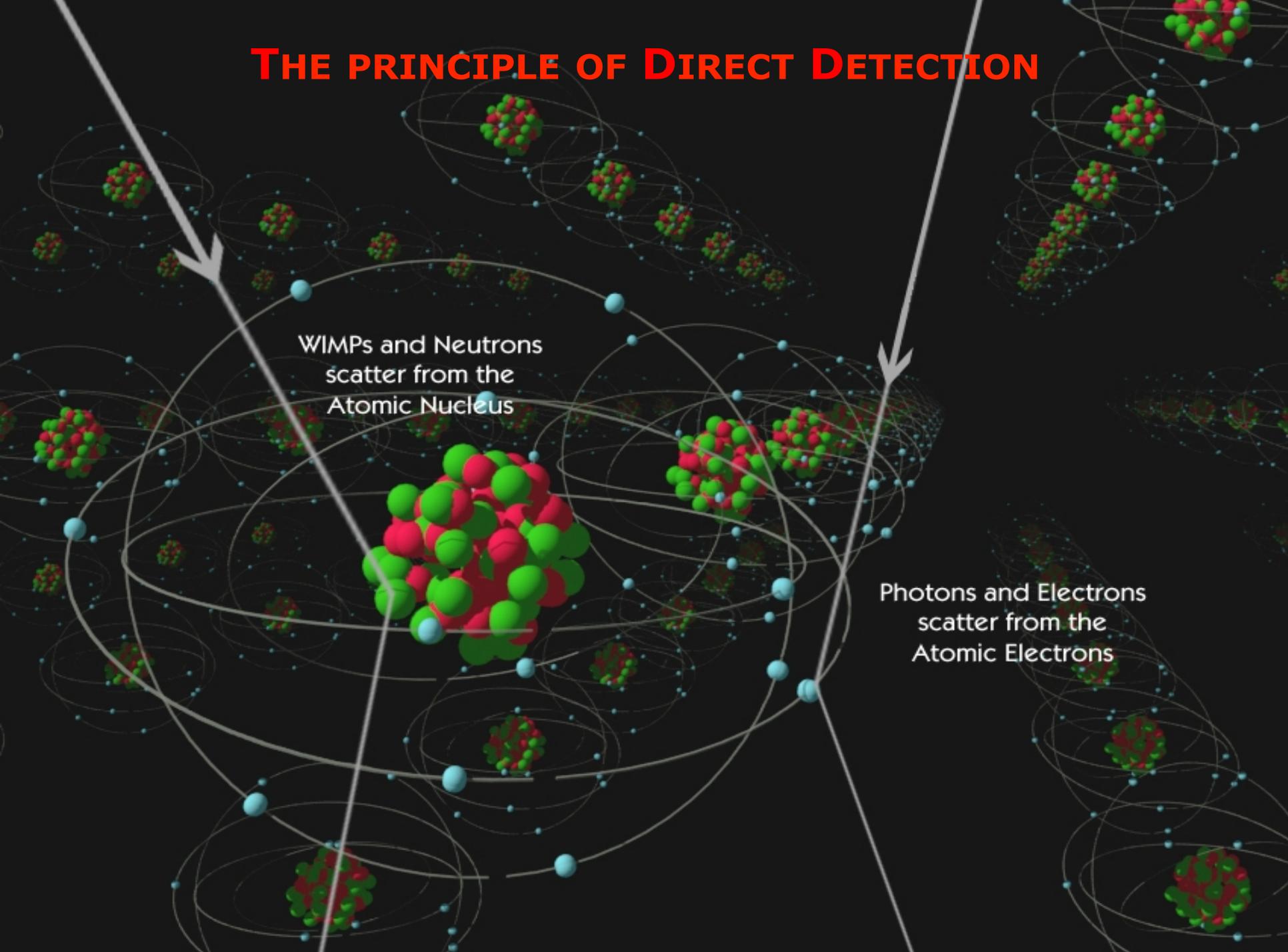
Scattering



Direct Detection



# THE PRINCIPLE OF DIRECT DETECTION



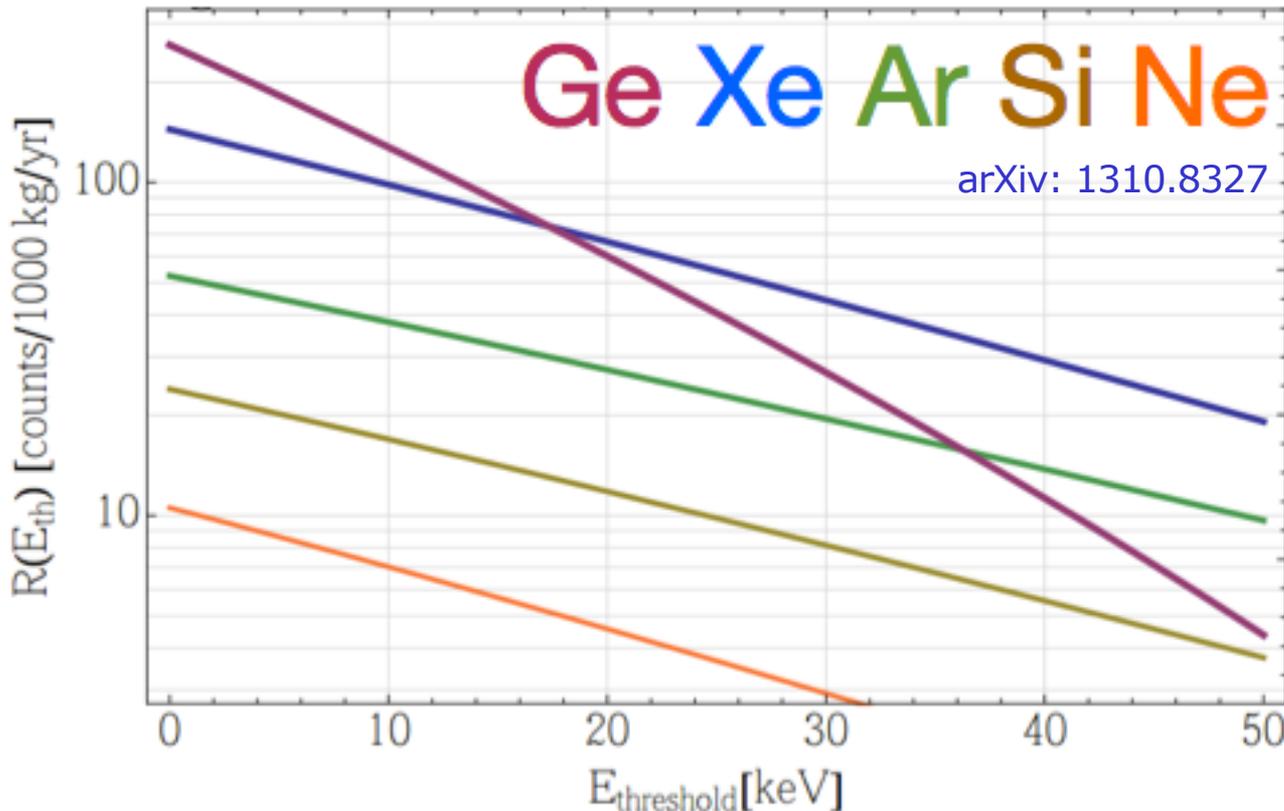
WIMPs and Neutrons  
scatter from the  
Atomic Nucleus

The diagram illustrates the principle of direct detection. It features a central atom with a nucleus made of red and green spheres and a cloud of blue electrons. Two white arrows point towards the atom from the top left and top right. The left arrow is labeled 'WIMPs and Neutrons scatter from the Atomic Nucleus' and the right arrow is labeled 'Photons and Electrons scatter from the Atomic Electrons'. The background is dark with several other atoms scattered around.

Photons and Electrons  
scatter from the  
Atomic Electrons

# NOBLE LIQUID DETECTORS FOR WIMPS

- ✓ Liquids provide a variety of targets (Xe, Ar, Ne)
- ✓ Easily scalable to ton-scale detectors + self-shielding
- ✓ Easy to purify for both electro-negative and radioactivity
- ✓ Large signals (high photon yield and high charge yield)
- ✓ Several background rejection techniques (S2/S1, PSD)



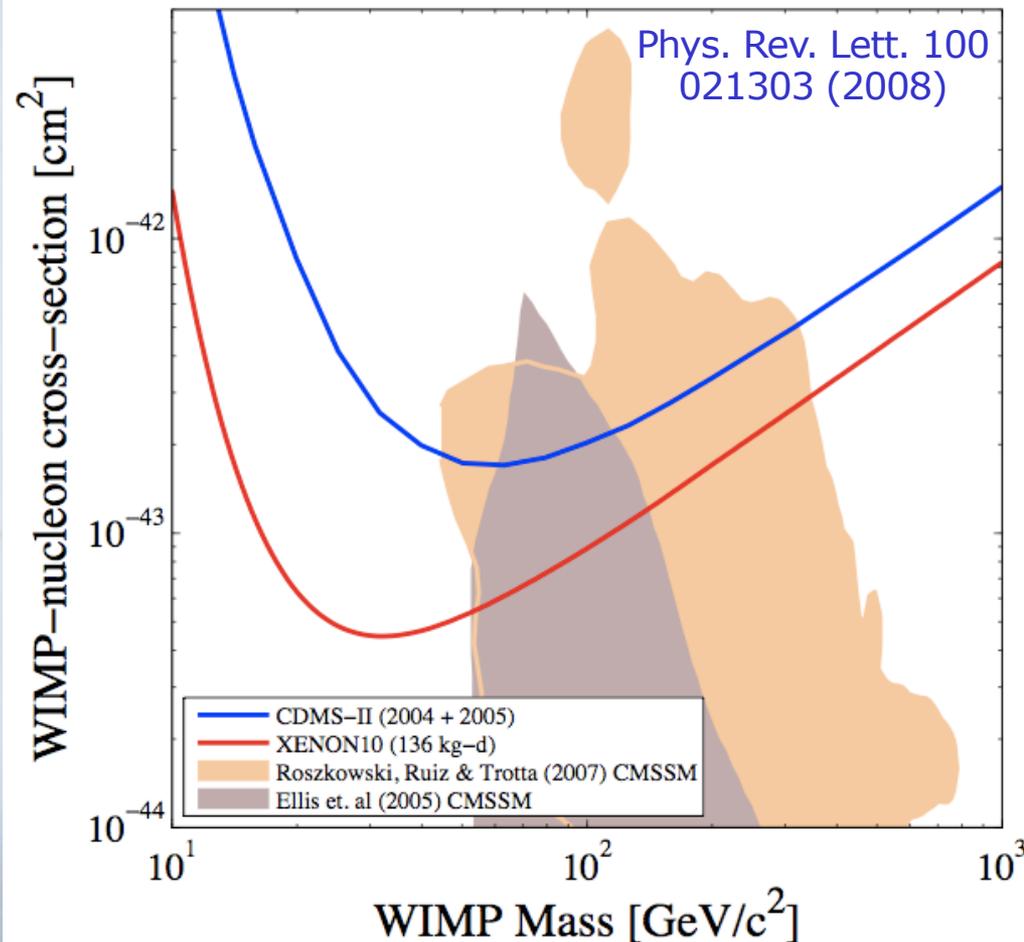
Example: WIMP interaction rates in five different DM targets

- $m_\chi$ : 100 GeV
- $\sigma_{\chi-n}$ :  $10^{-45}$  cm<sup>2</sup>
- 1 ton·year

# THE XENON-10 EXPERIMENT AT GRAN SASSO

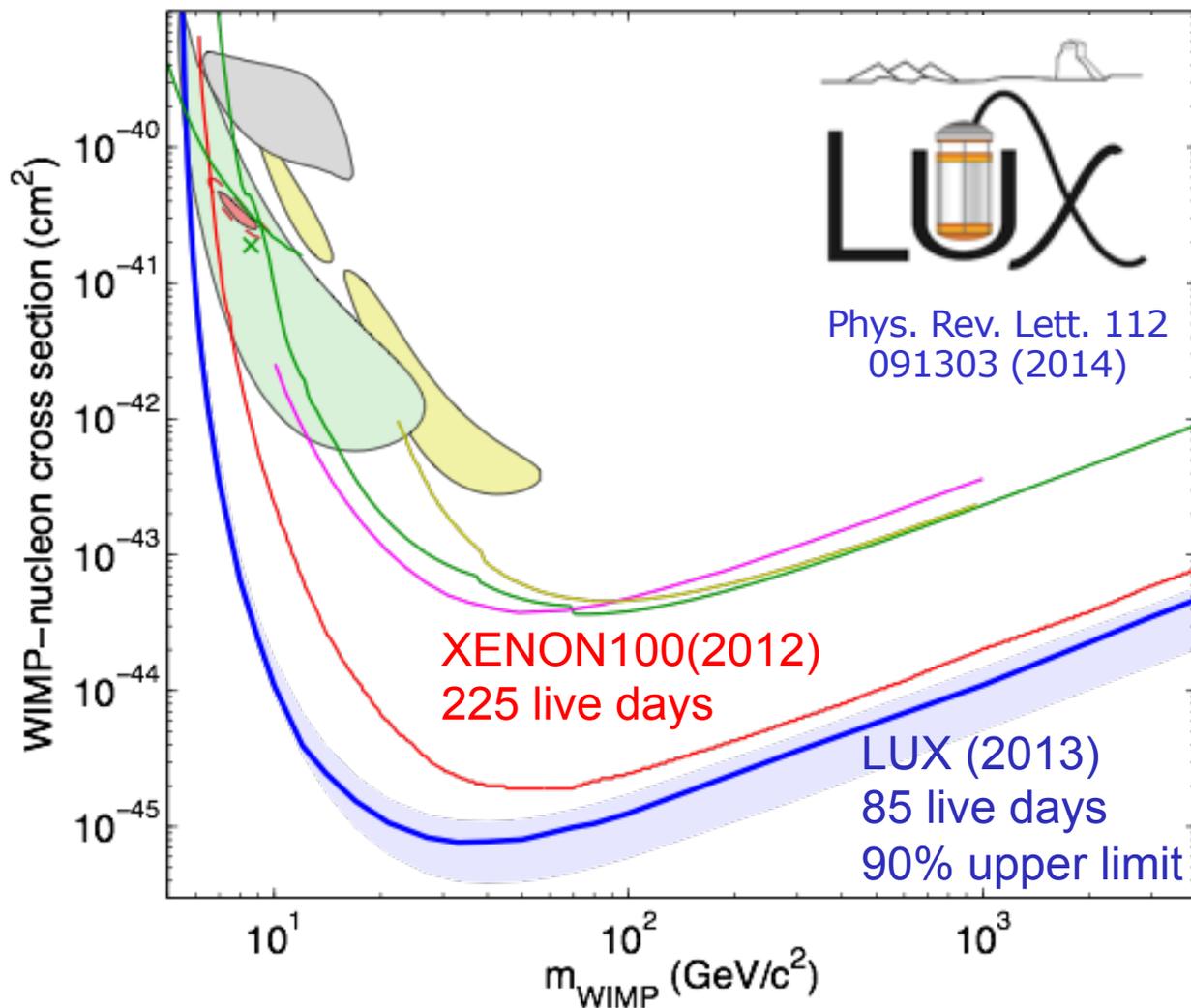
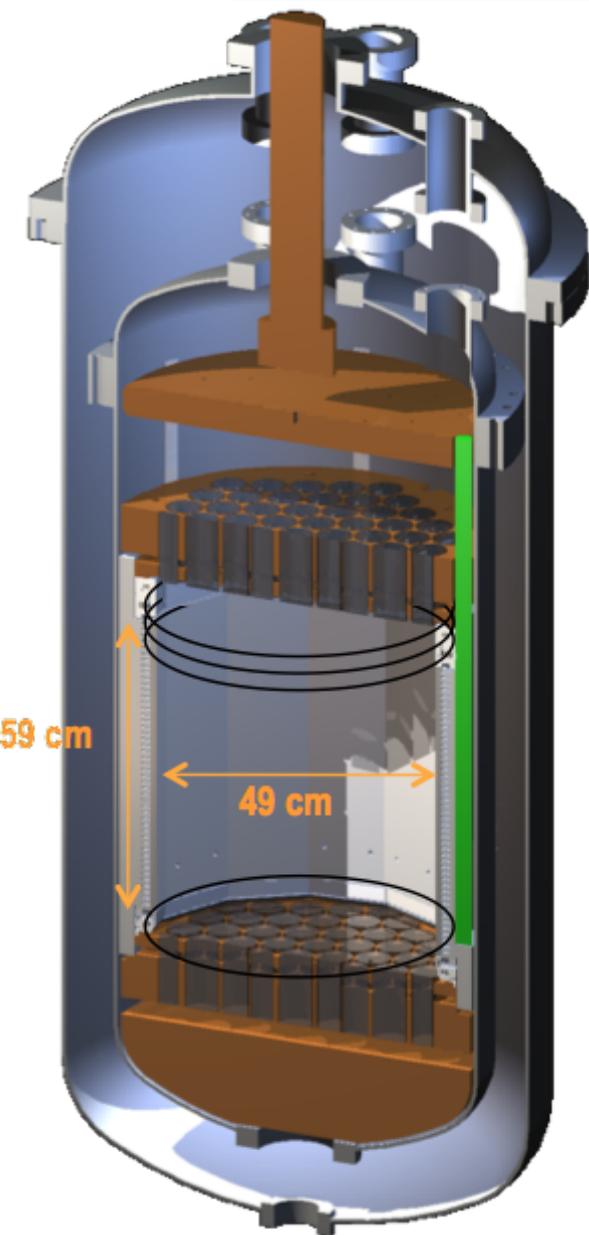


- Operated in 2006-2007
- 15 kg Liquid Xenon target
- $\sigma < 4.5 \cdot 10^{-44} \text{ cm}^2 @ 30 \text{ GeV}$

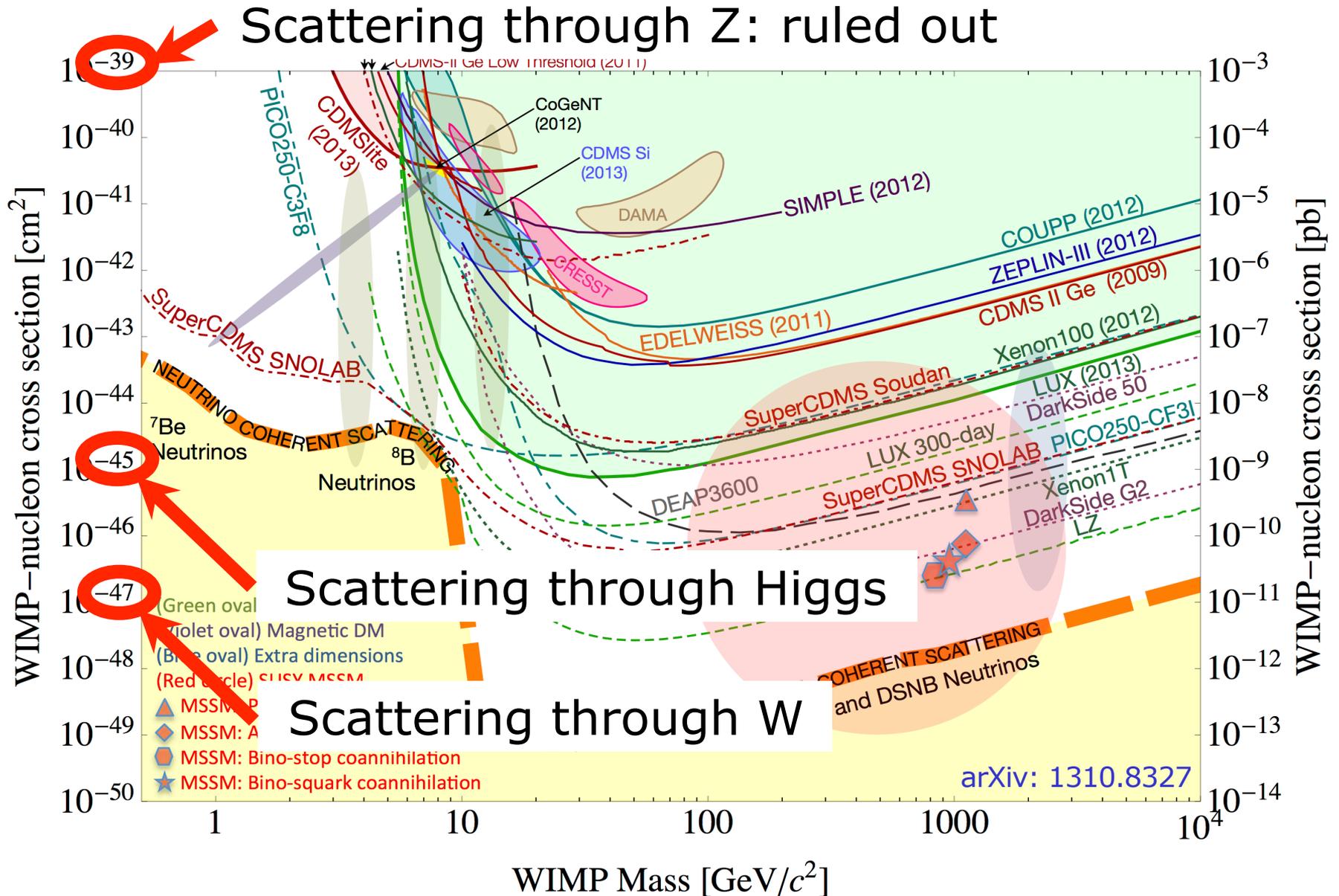


# THE LARGE UNDERGROUND XENON EXPERIMENT

- world's most sensitive experiment
- $\sigma < 7.6 \cdot 10^{-46} \text{ cm}^2$  @33 GeV (2013)



# STATE OF THE ART IN DIRECT DETECTION



# THE DARKSIDE PROGRAM AT LNGS

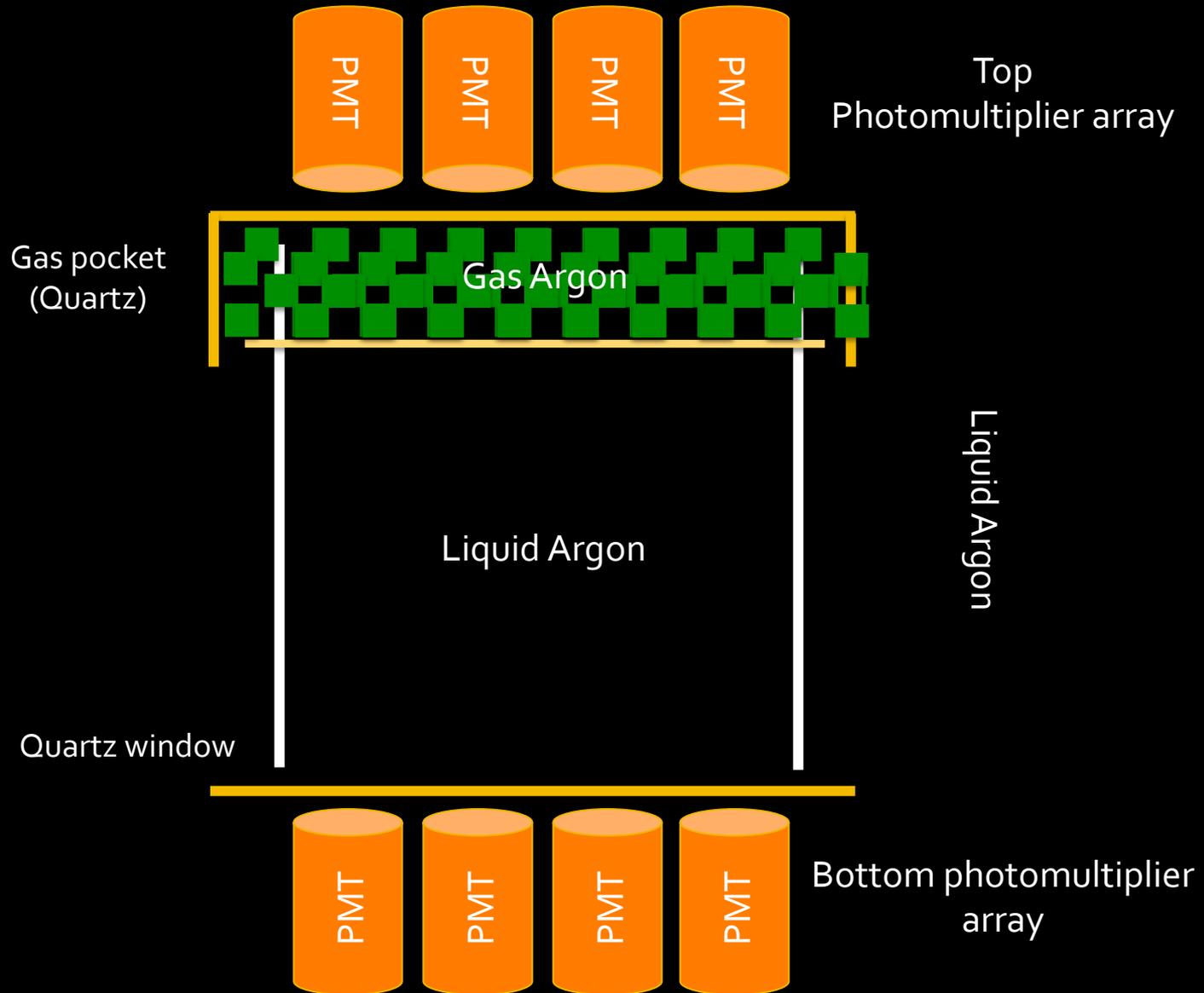


- DS-50: took data with atmospheric Argon in 2014 (presented here)
- sensitivity goal  $10^{-45}$  cm<sup>2</sup>

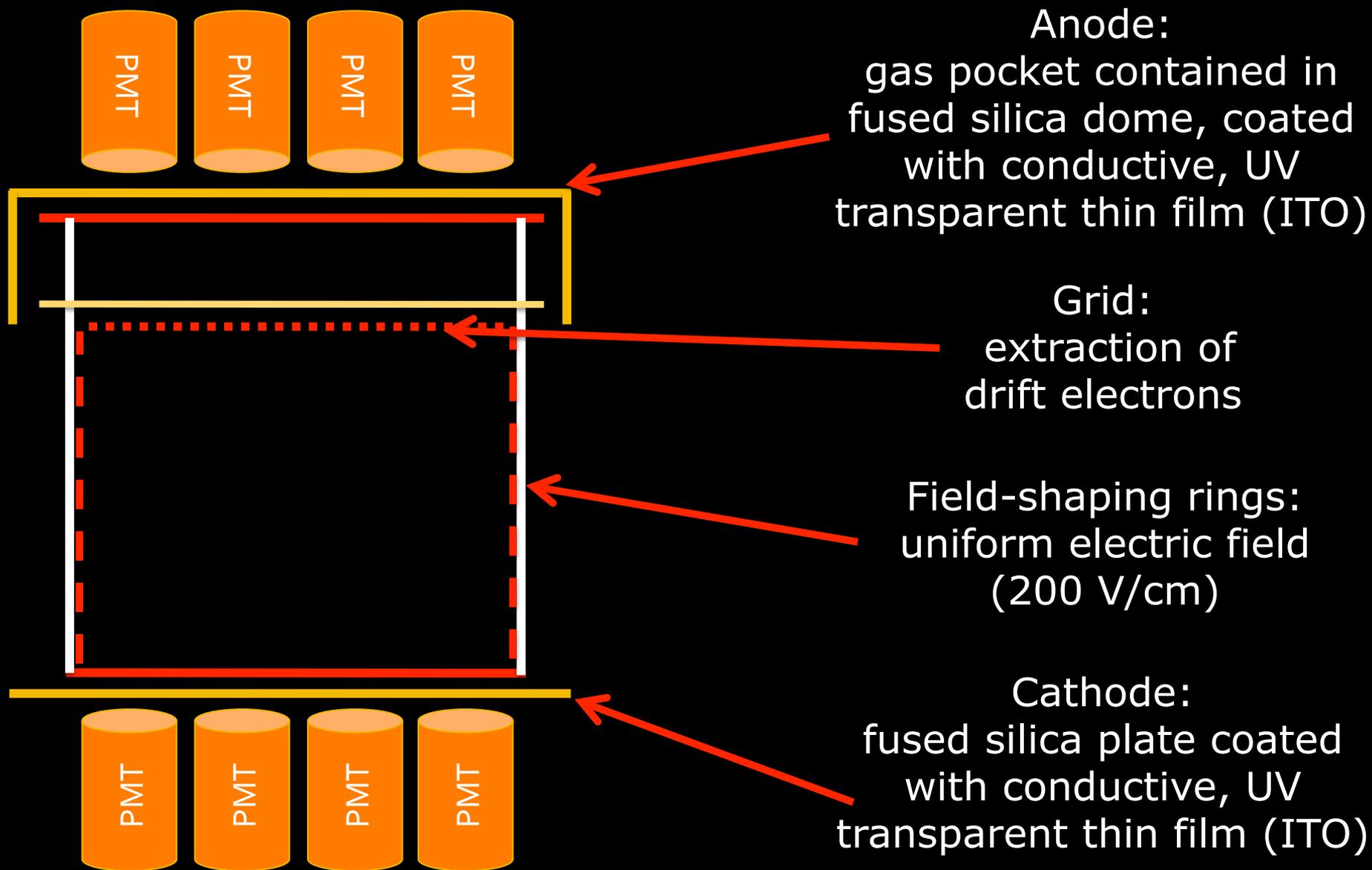
- Dual-phase argon TPC
- Underground argon, depleted in <sup>39</sup>Ar
- Three-fold discrimination (S1 pulse shape, S2/S1, 3D reconstruction)
- TPC immersed in active neutron and muon vetoes



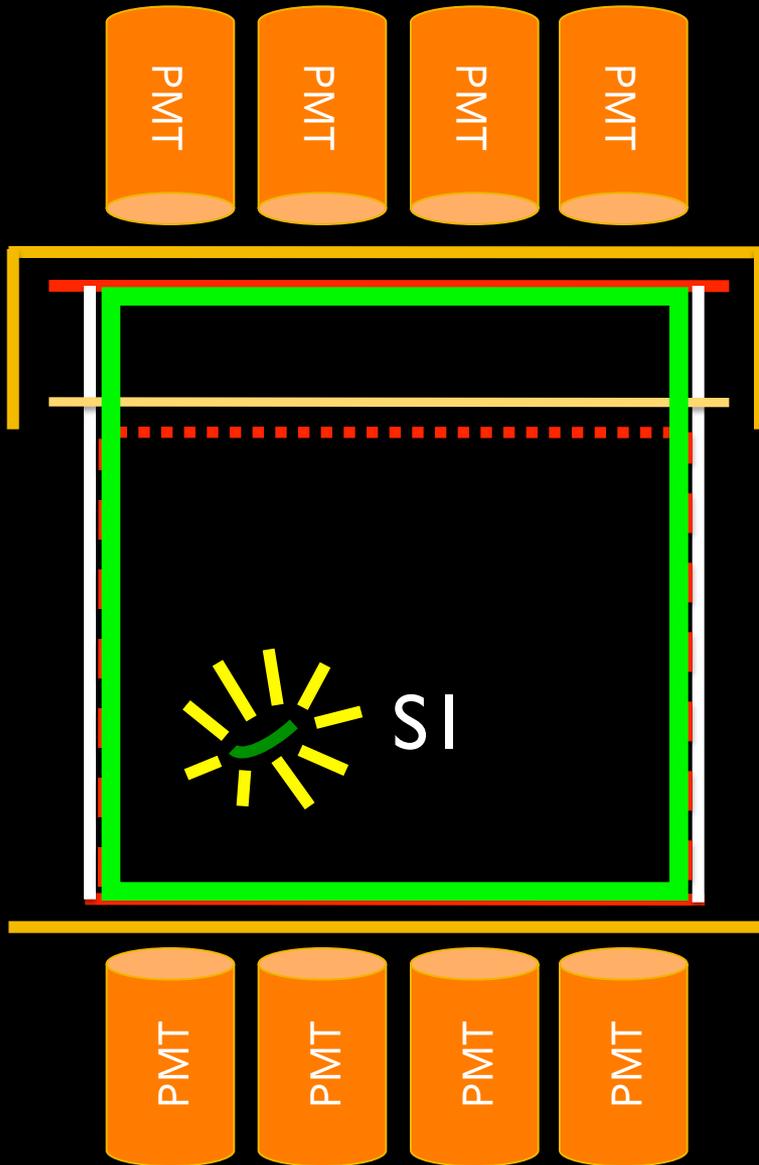
# DUAL-PHASE NOBLE LIQUID TPC



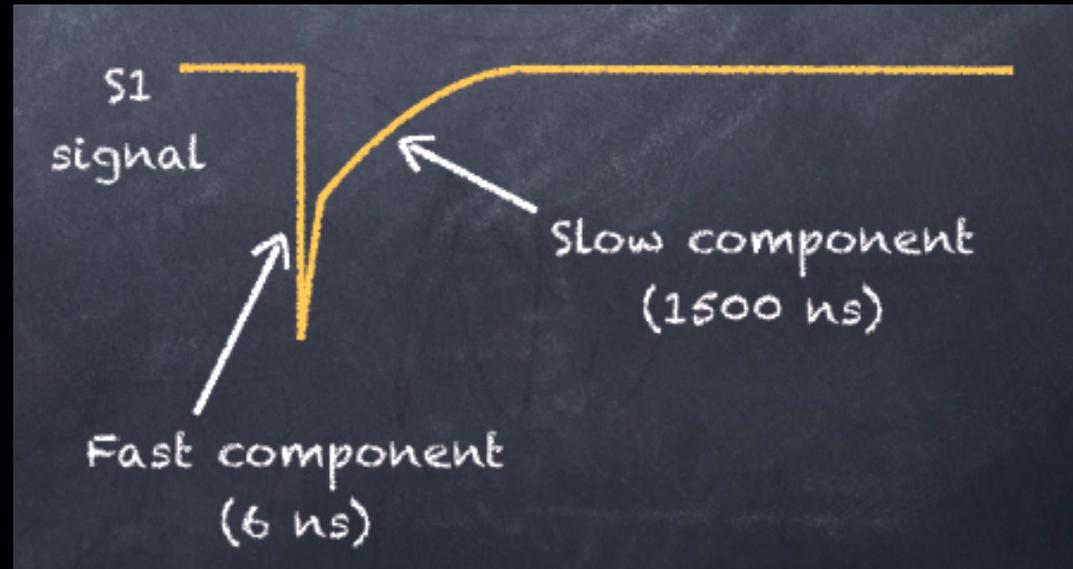
# DUAL-PHASE NOBLE LIQUID TPC



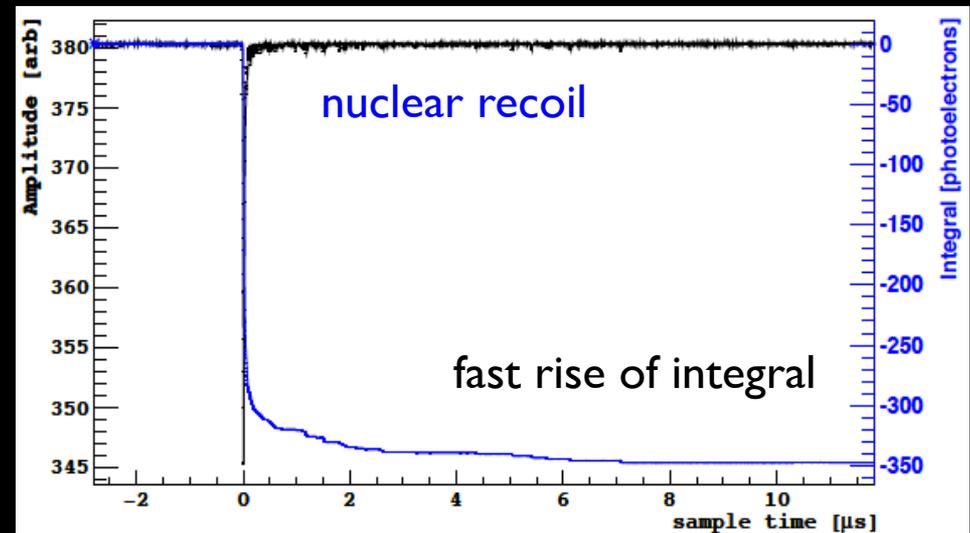
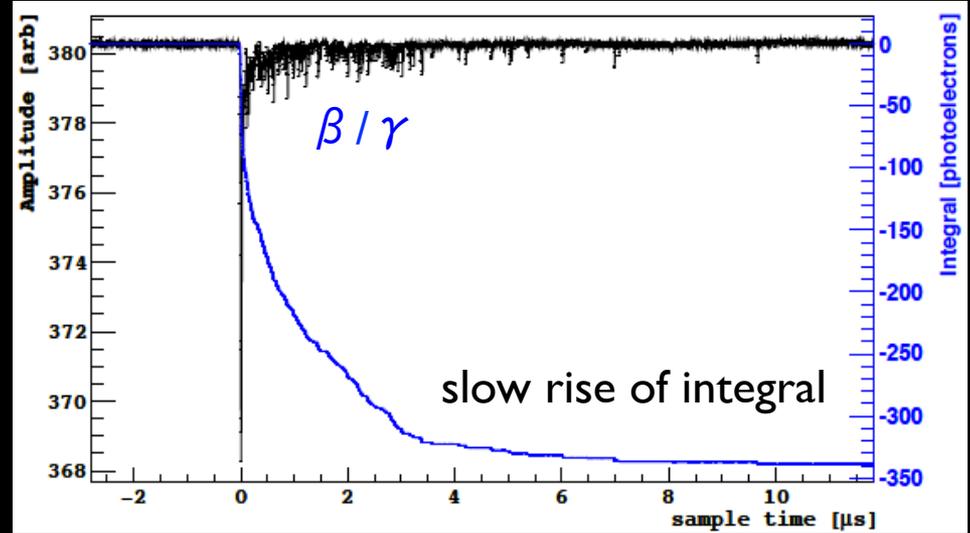
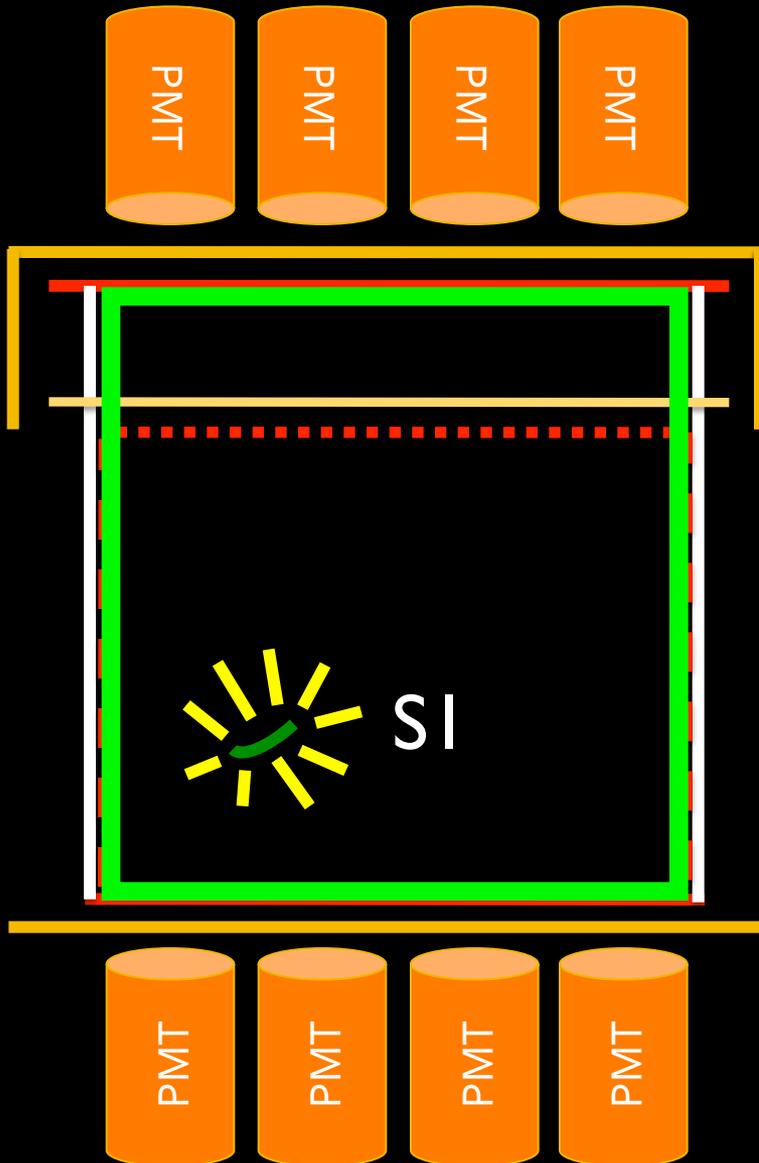
# S1 SIGNAL SHAPE IN ARGON



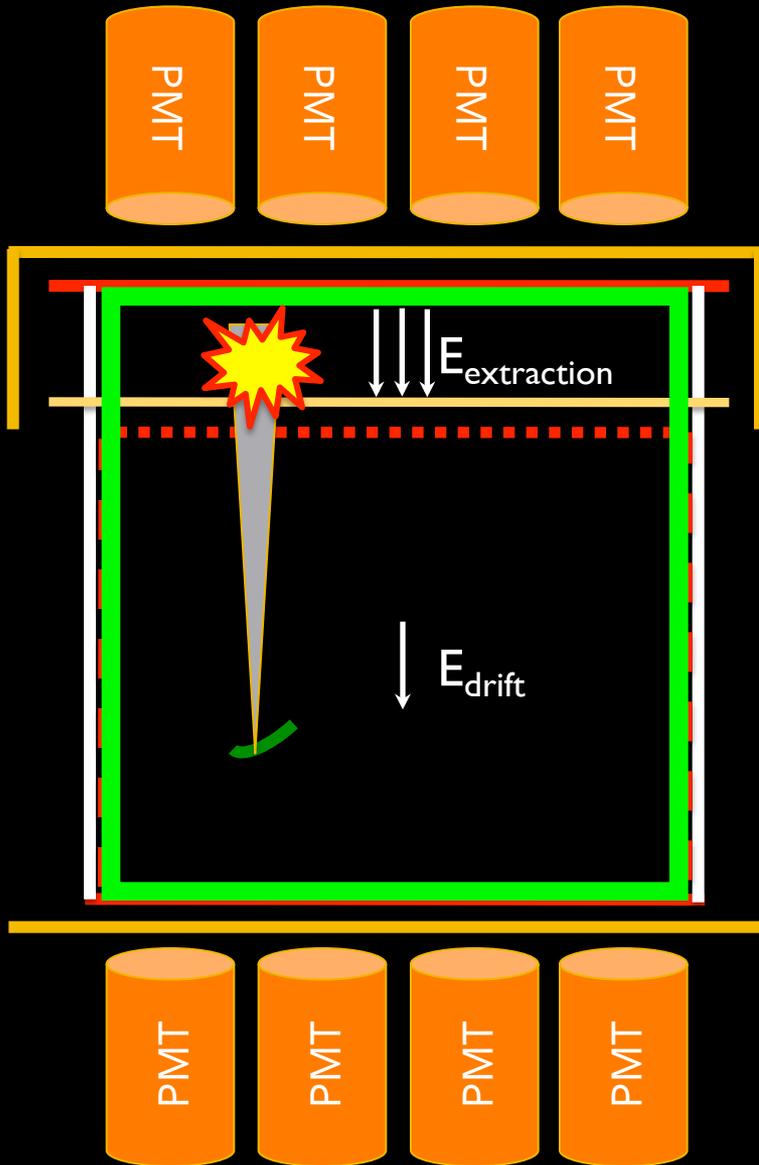
Nuclear and electron recoils produce ionization and excitation along their tracks. Excited  $\text{Ar}_2^*$  are formed and their de-excitation leads to the emission of scintillation light with two separate components, fast and slow (associated to  $\text{Ar}_2^*$  singlet and triplet state). The distribution of light on the two components is very strongly dependent on  $dE/dx$ .



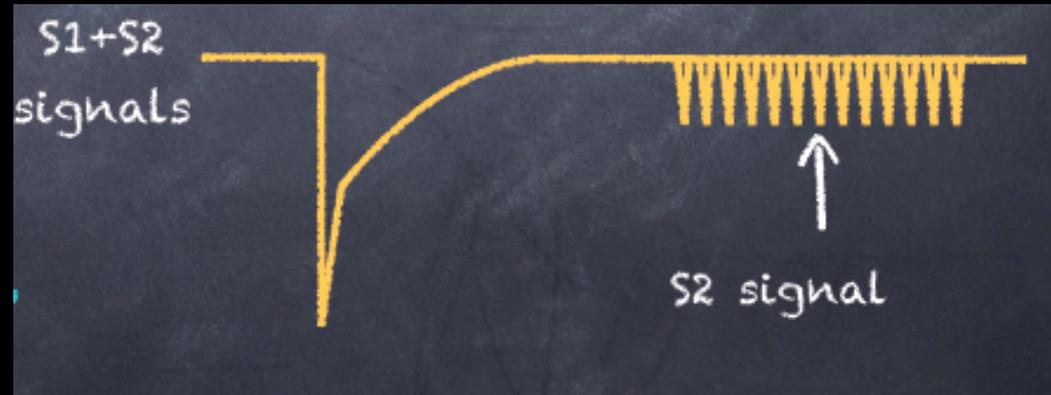
# S1 PULSE SHAPE DISCRIMINATION



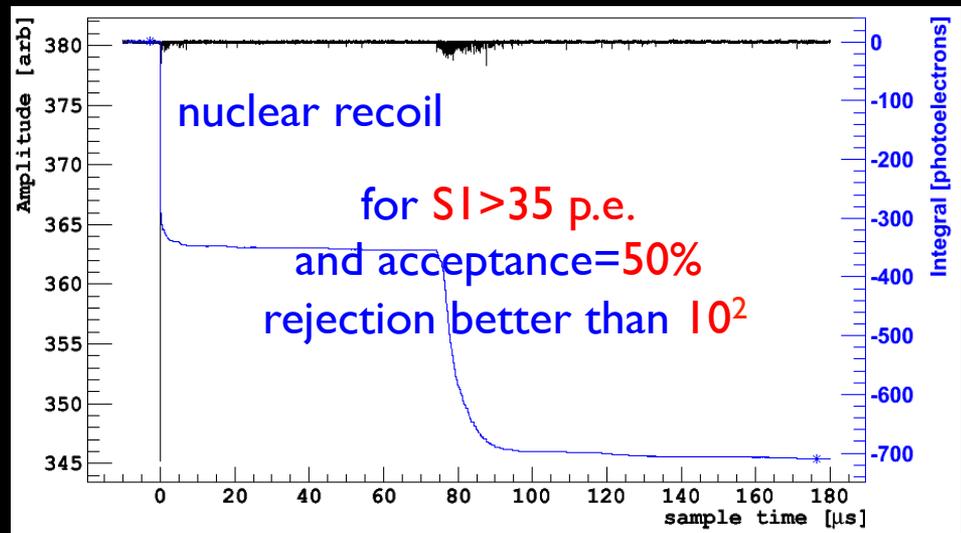
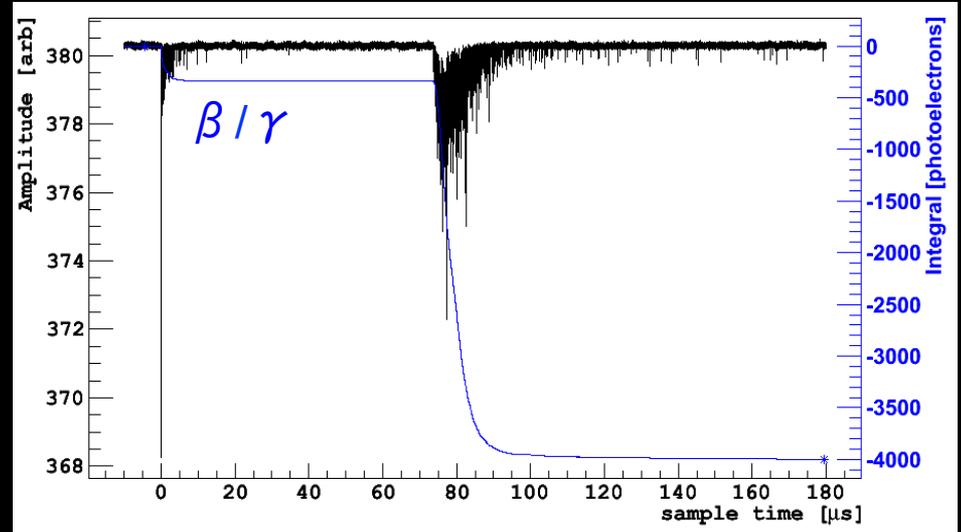
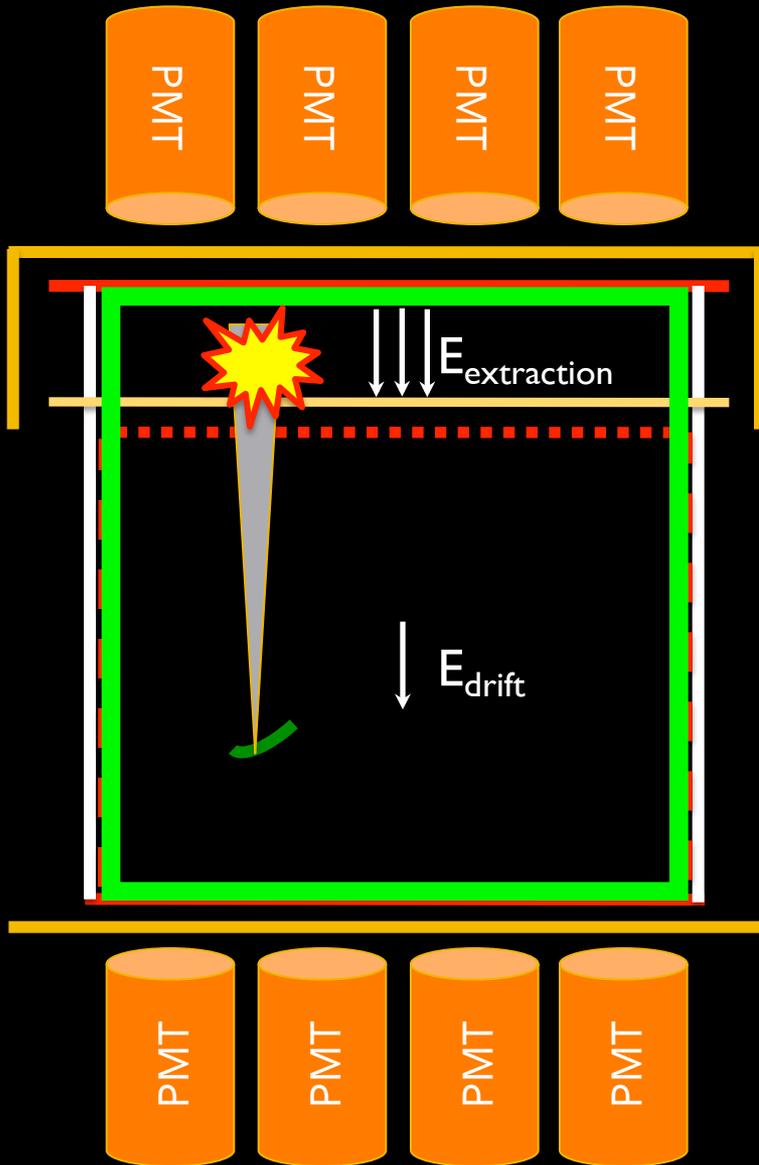
# S2 SIGNAL GENERATION



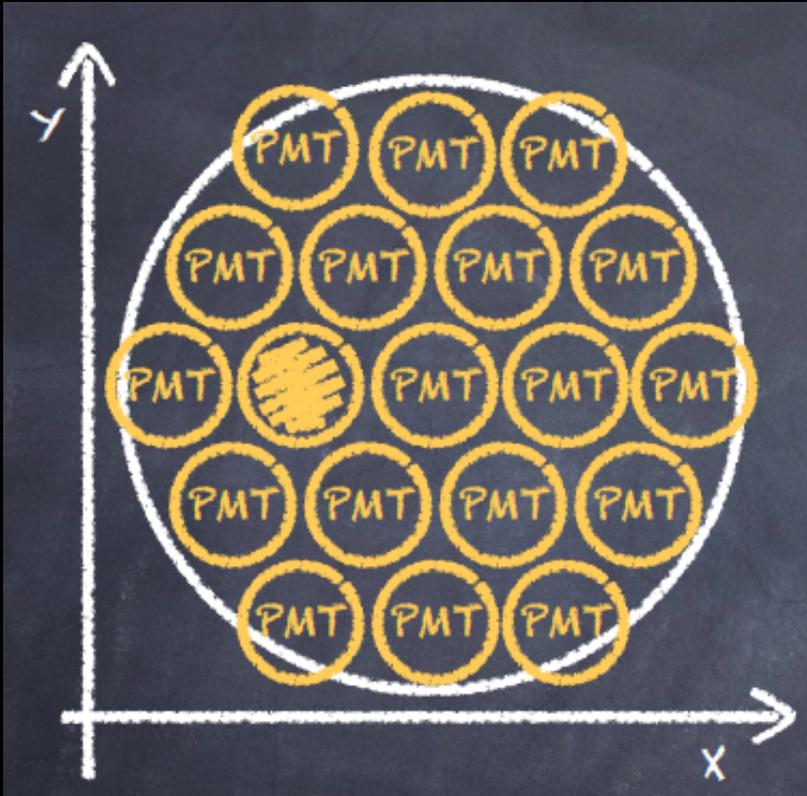
The ionization electrons surviving the recombination are drifted towards the liquid-gas interface ( $E_{\text{drift}}=200$  V/cm). A field of 2.8 kV/cm is applied to fully extract electrons to the gas phase. Electroluminescence in gas produces a secondary scintillation signal (S2) proportional to the ionization signal. The ratio  $S2/S1$  depends very strongly on the probability of surviving the initial recombination, and therefore on  $dE/dx$ .



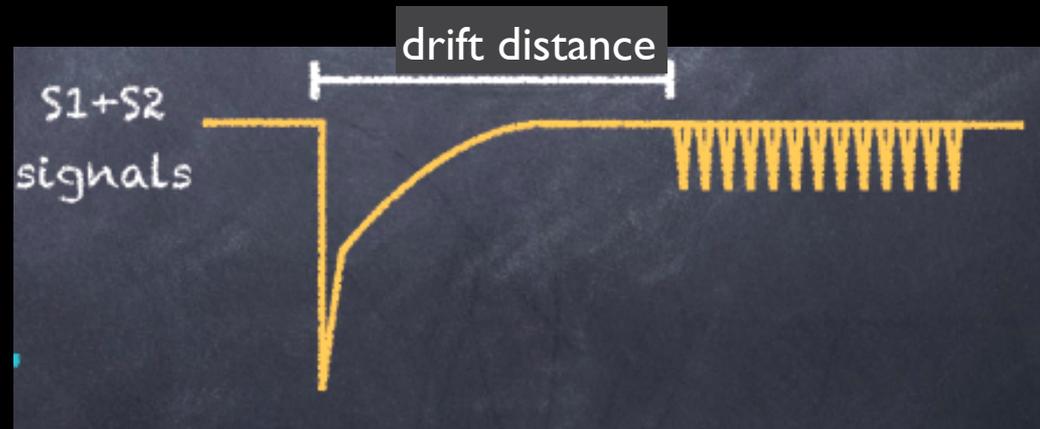
# S2/S1 DISCRIMINATION



# 3D EVENT LOCALIZATION



The time difference between **S1** and **S2** corresponds to the  **$e^-$  drift time** and it is used to estimate the z coordinate of the interaction. Drift speed in Argon at **200 V/cm** is about **1 mm/ $\mu$ s**.

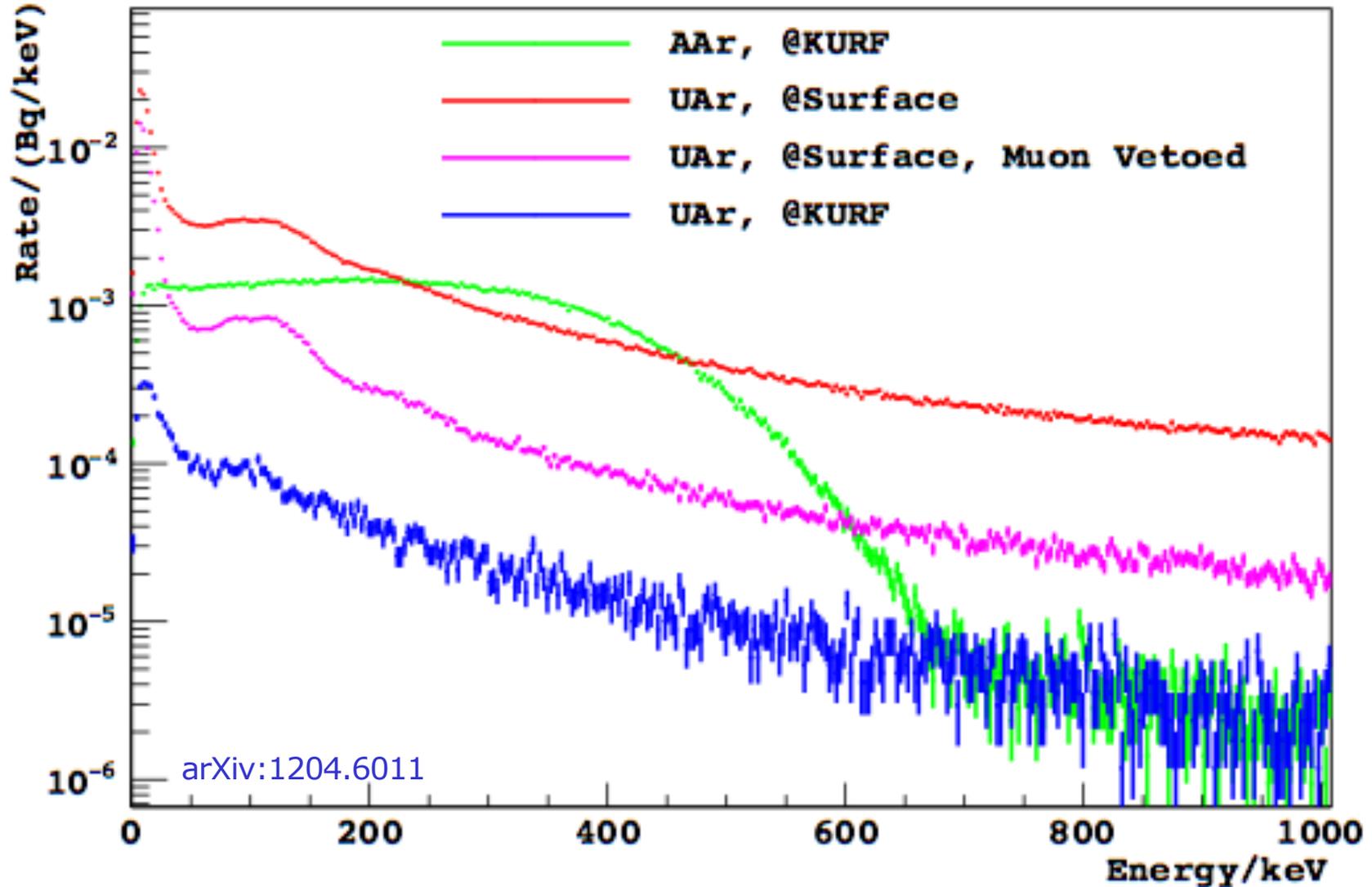


Since **S2** production occurs very close to the top PMTs, the signal distribution will be strongly not uniform and can be used to locate the ionization event in the **x-y plane**

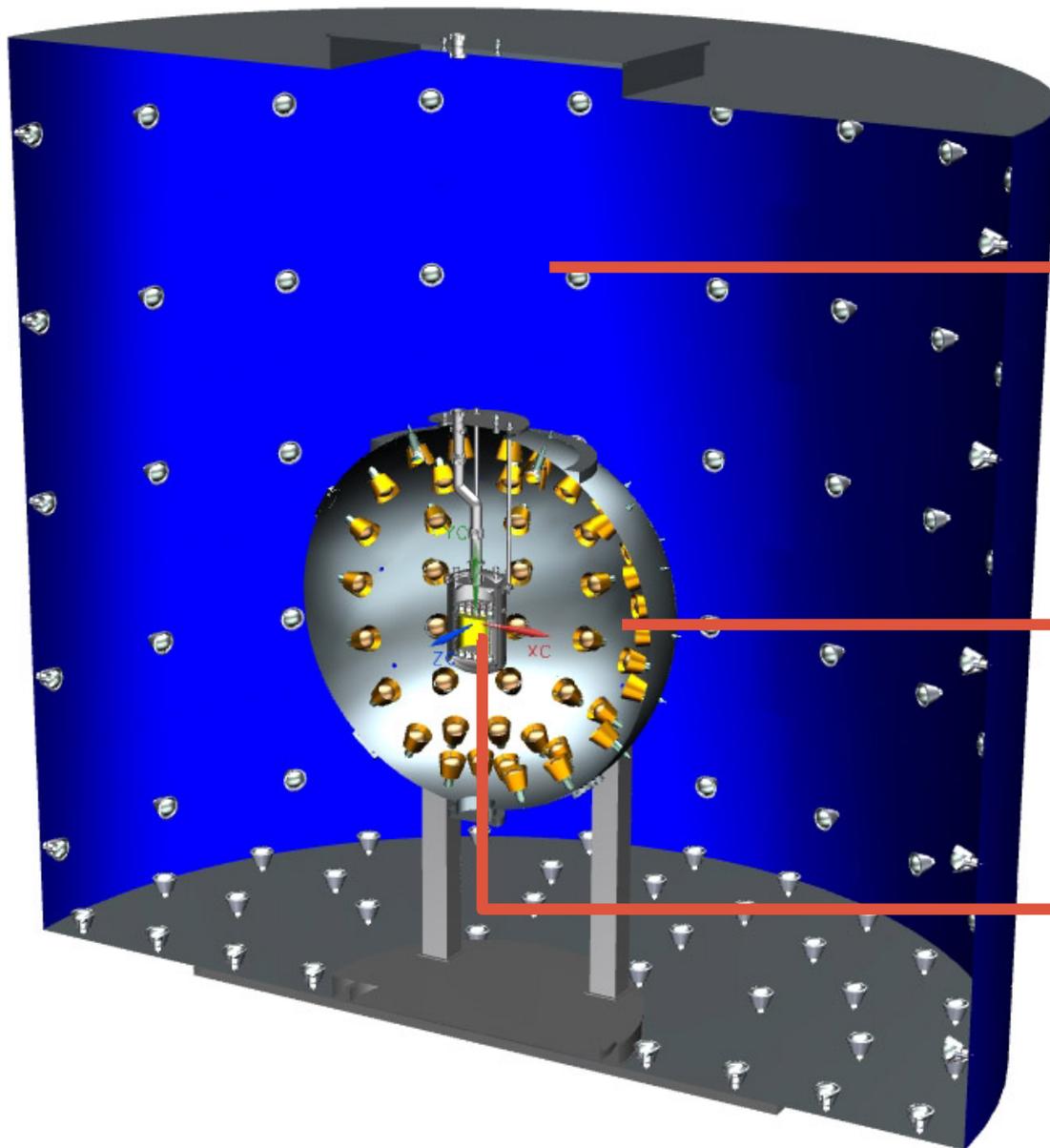
3D localization allows the implementation of fiducial volume cuts (thereby removing external background)

# UNDERGROUND ARGON: DEPLETION FACTOR > 100

## Underground Argon Measurements



# DARKSIDE-50 WITHIN ITS VETO DETECTORS



## Water Čerenkov detector:

- active muon veto +
- passive neutron veto
- 11 m wide, 10 m tall
- 80 PMTs in Water Tank

## Liquid Scintillator detector:

- active neutron veto
- boron-loaded scintillator
- 4 m diameter (20ns)
- 110 PMTs in Scintillator

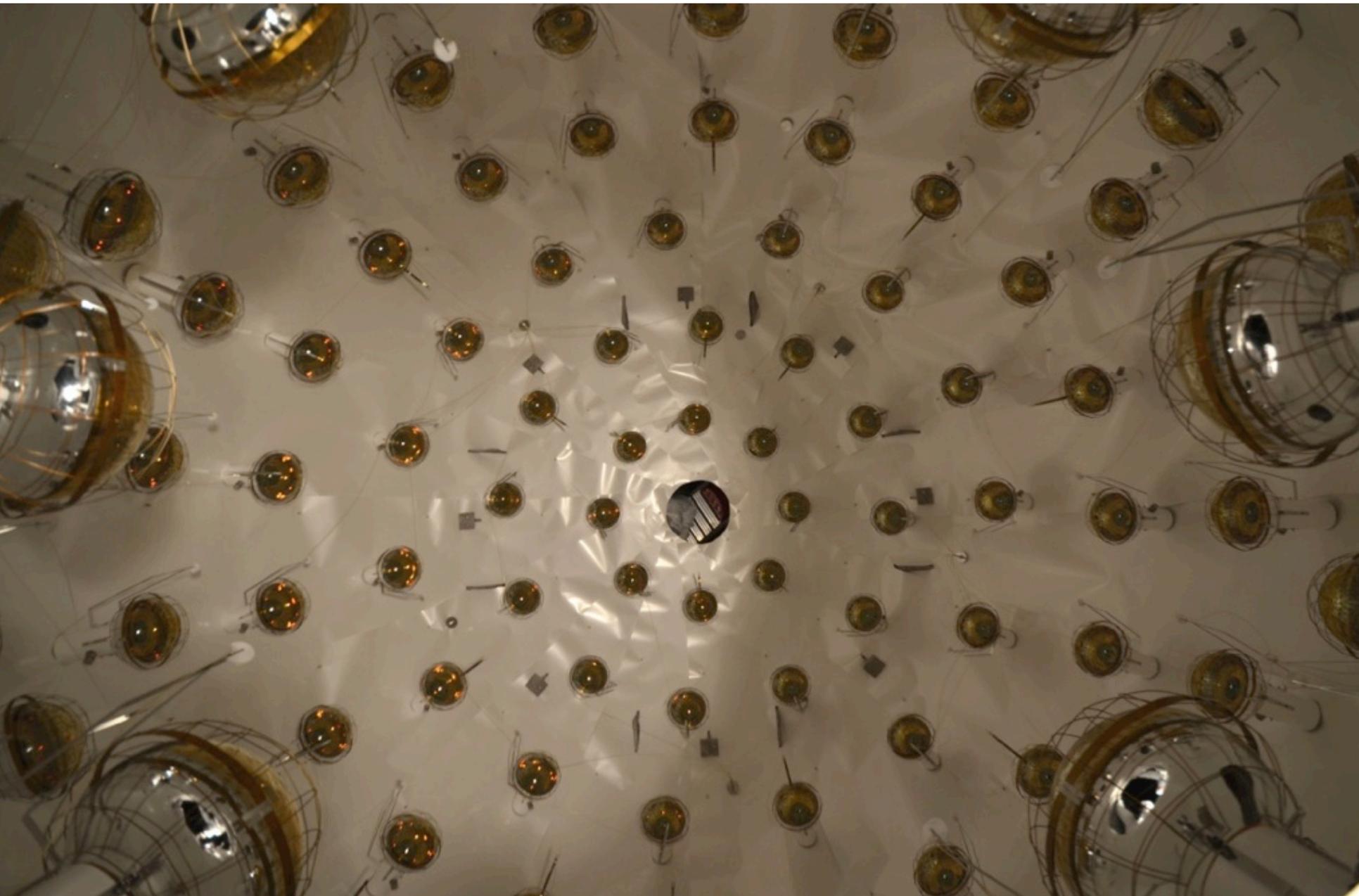
## Liquid Argon TPC

- 50 kg active mass
- 2x19 PMT arrays
- 37 kg fiducial volume

# TPC CONSTRUCTION – SUMMER 2013



# NEUTRON VETO DETECTOR DURING INSTALLATION

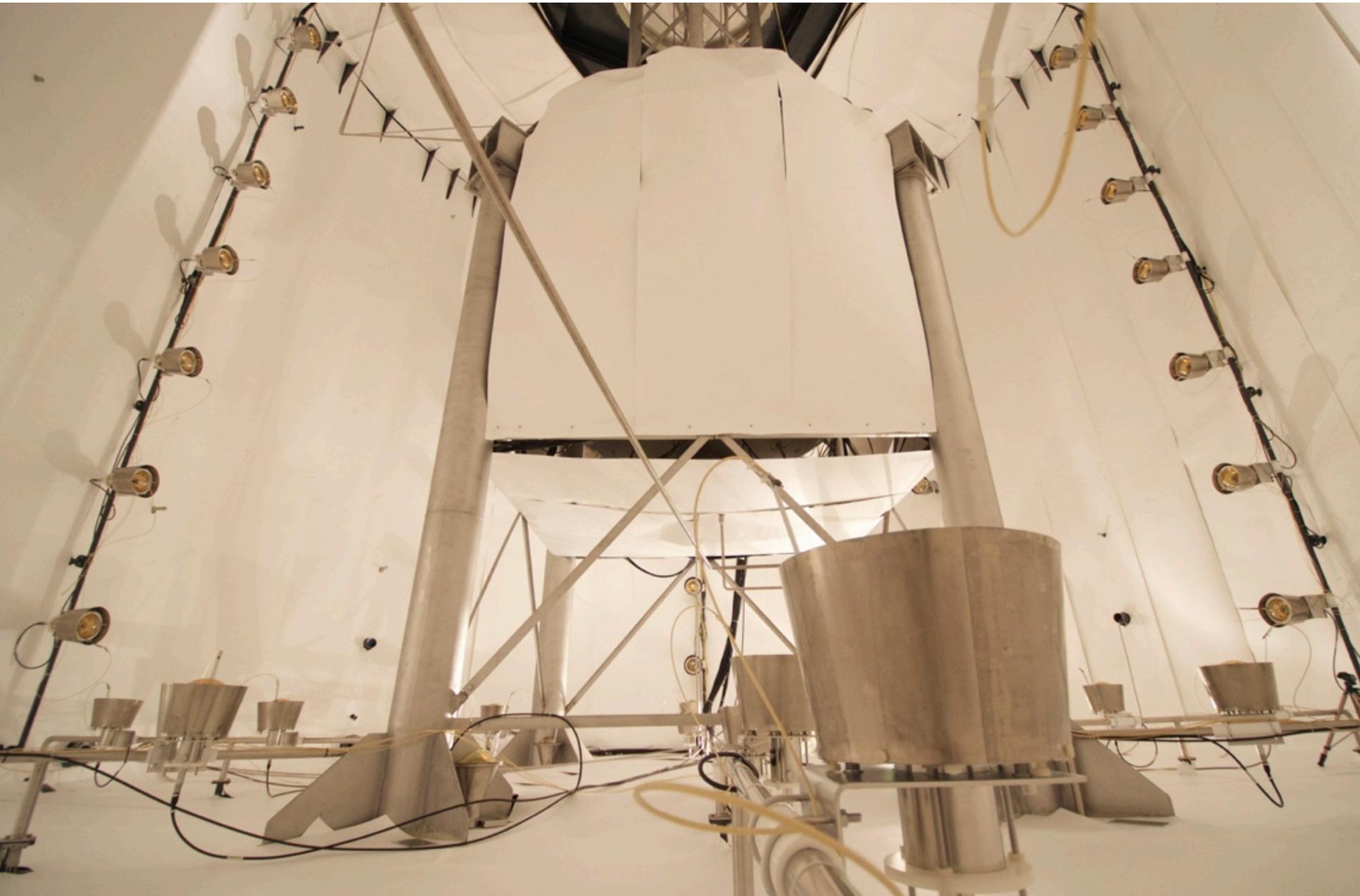




# NEUTRON VETO WITH TPC CRYOSTAT



# MUON VETO, RIGHT BEFORE WATER FILLING

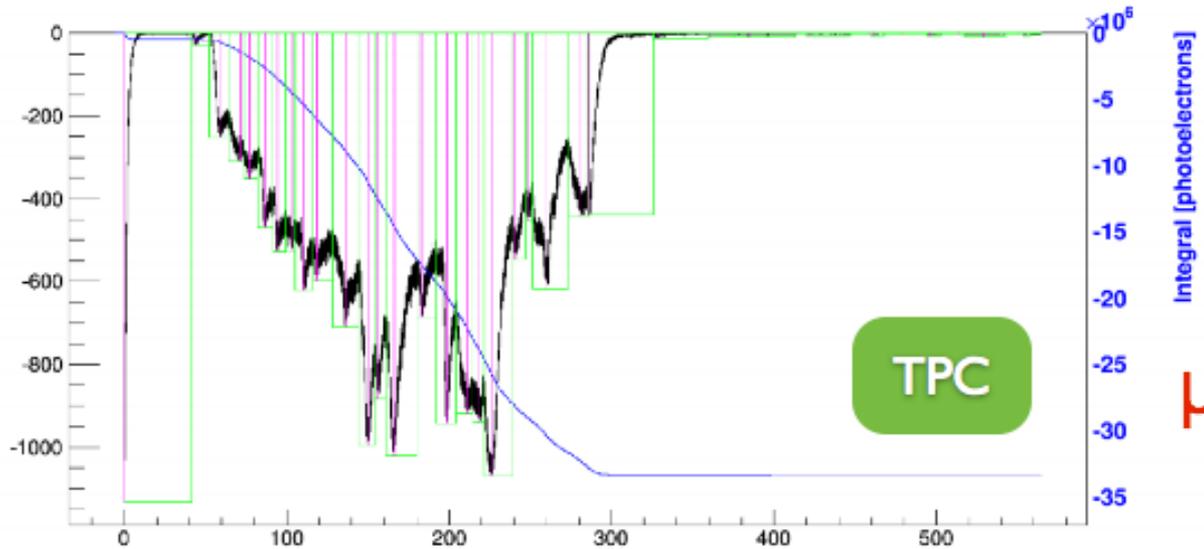


# DARKSIDE-50 TIMELINE

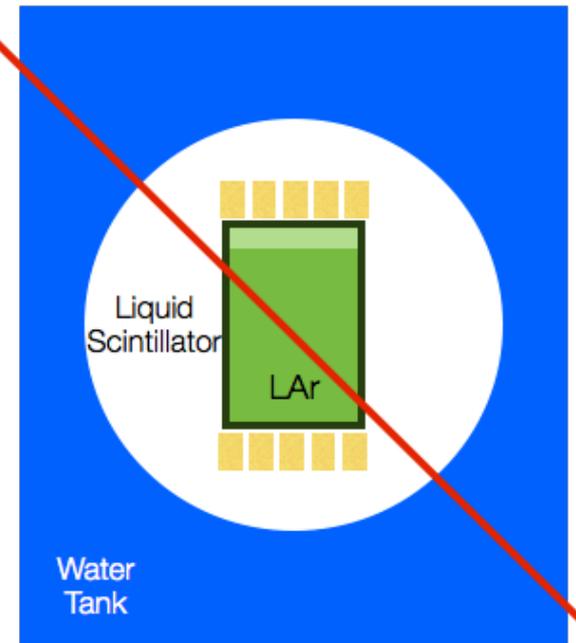
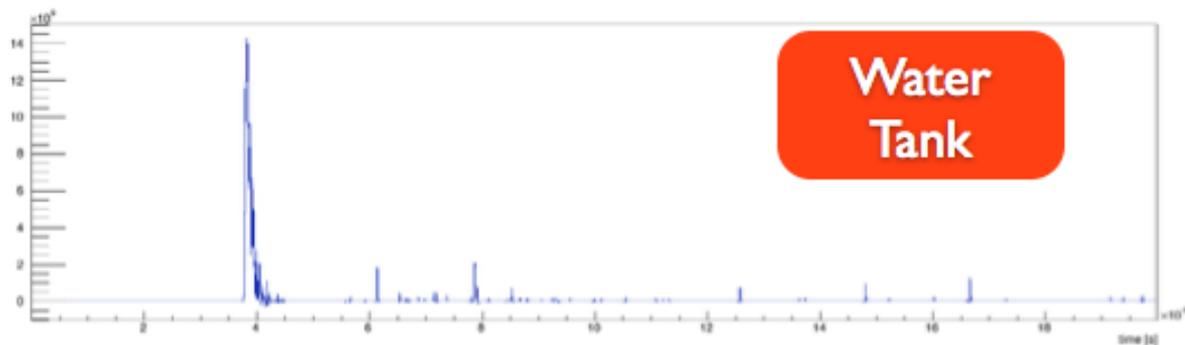
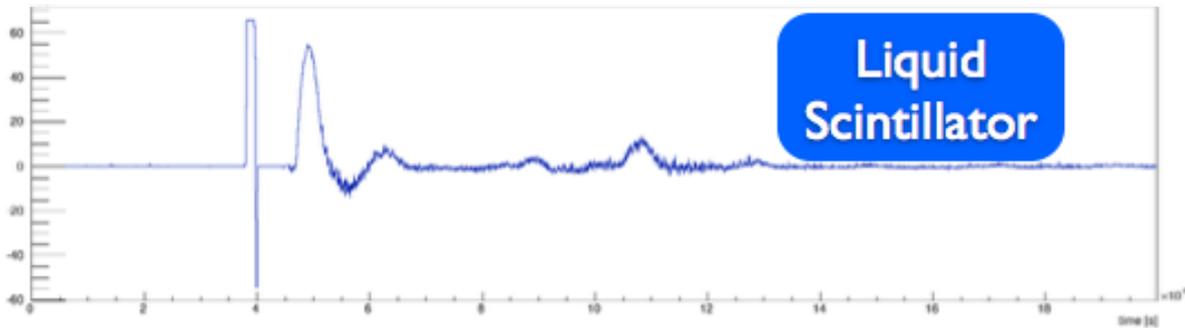
---

- **Nov. 2013:** All 3 detectors filled
  - Coordinated data taking started Nov. 6
  - TPC filled with atmospheric argon (AAr)
  - LSV filled with high-radioactivity TMB
- **Nov. - Jan. 2014:** mainly DAQ/trigger improvements
  - accumulated 6 live-days (exposure 280 kg·day)
  - equivalent to  $\sim 3$  years with underground argon
  - PSD performance presented at UCLA-DM 2014
- **Nov. - June 2014:** dataset discussed here
  - High  $^{14}\text{C}$  TMB: degraded veto performance
  - 47.1 live days exposure (1422 kg·day fiducial)
  - $^{39}\text{Ar}$  statistics corresponds to 19.4 years of UAr
- **Jun. 2014 – Feb. 2015:** veto scintillator operations
- **Oct. 2014 – Feb. 2015:** source calibration campaign

# ALL SYSTEMS CHECK OUT!

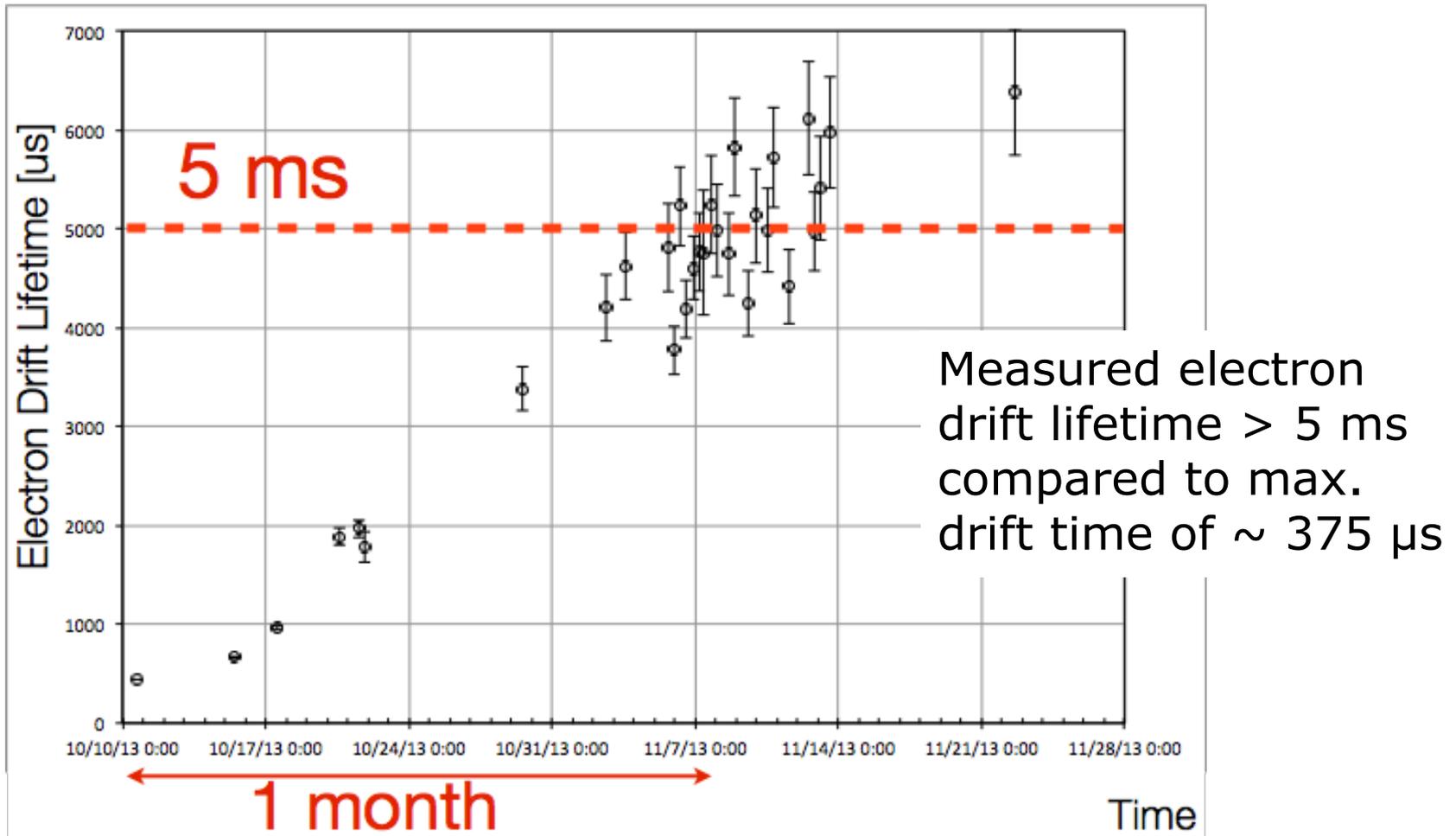


$\mu$



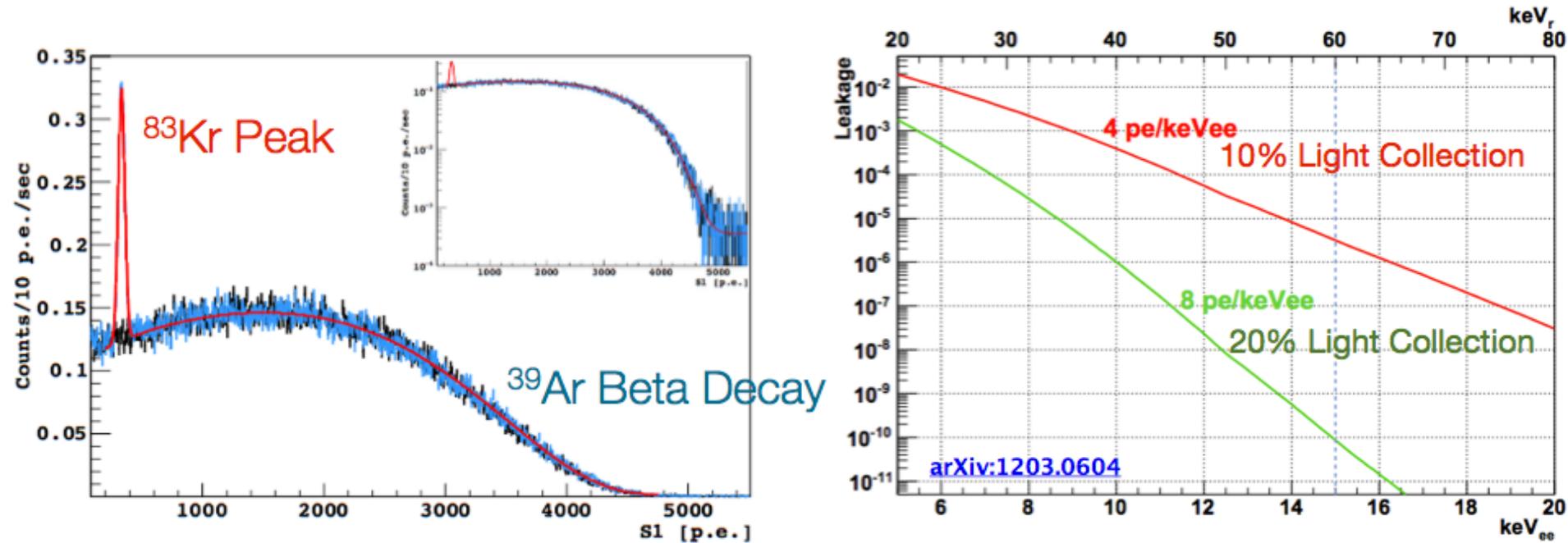
# TPC COMMISSIONING: PURITY

- Closed loop argon recirculation ( $\sim 30$  slpm)
- Gaseous phase purification using commercial getter
- Cryogenic charcoal trap to remove Rn contamination



# TPC COMMISSIONING: LIGHT YIELD

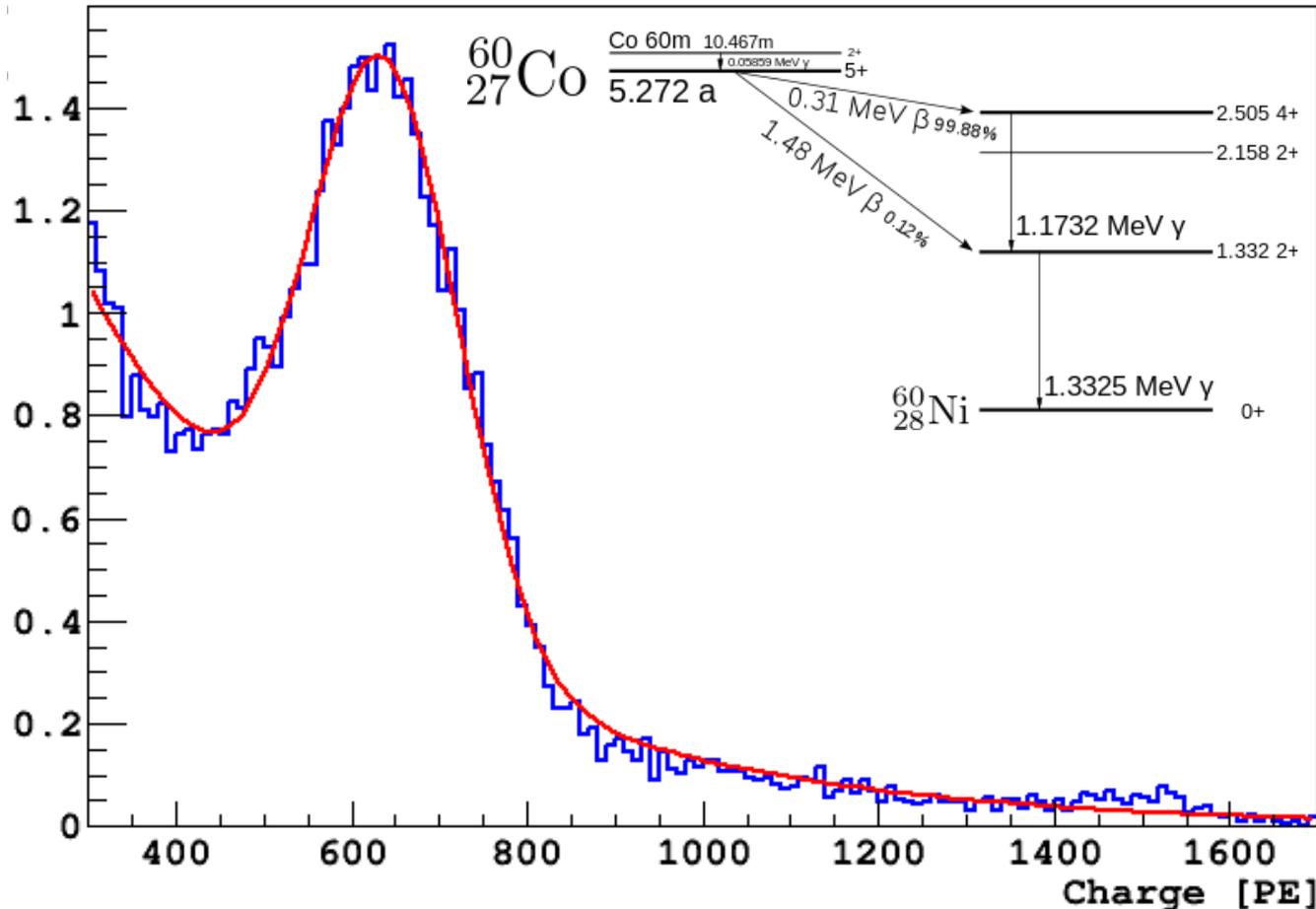
- TPC filled with atmospheric argon (1 Bq/kg)
- $^{83m}\text{Kr}$  gas deployed into detector ( $41.5 \text{ keV}_{ee}$ )



- Light yield at zero field:  $(7.9 \pm 0.4) \text{ PE/keV}$
- Light yield at 200 V/cm:  $(7.0 \pm 0.3) \text{ PE/keV}$
- Exceeding design requirements of 6.0 PE/keV

# NEUTRON VETO COMMISSIONING: LIGHT YIELD

- Neutron veto setup to trigger on events in the LAr TPC
- High energy coincident  $^{60}\text{Co}$  events from cryostat steel
- LY was measured by combining  $^{60}\text{Co}$  and  $^{14}\text{C}$  spectra
- $\text{LY} = (0.54 \pm 0.04) \text{ PE/keV}$  (exceeding design specs.)



- High  $^{14}\text{C}$  rate in (biogenic) TMB
- **$\sim 98\%$  LSV efficiency in this dataset**
- TMB replaced with low- $^{14}\text{C}$  (fossil) batch
- $^{14}\text{C}$  content measured at LLNL via AMS prior to filling

# ANALYSIS CUTS: DAQ & QUALITY

	Cut	Residual Livetime
Run	Usable runs	$(53.8 \pm 0.2)$ d
	Automated selection	$(51.1 \pm 0.2)$ d
	Single run	$(48.8 \pm 0.2)$ d
Quality	Baseline found	$(48.8 \pm 0.2)$ d
	Time since previous trigger	$(48.7 \pm 0.2)$ d
	Large gap	$(48.1 \pm 0.2)$ d
	Veto data present	$(47.1 \pm 0.2)$ d
	Total	$(47.1 \pm 0.2)$ d

9.3% of livetime removed  
by run-level quality cuts

3.5% of livetime removed  
by event-level quality cuts

Uncertainty dominated by the accuracy of the trigger board

# ANALYSIS CUTS: PHYSICS CUTS

	Cut	Acceptance	Fiducial Mass
Physics	Number of pulses	$0.95^{+0.00}_{-0.01}$	
	First pulse time	$1.00^{+0.00}_{-0.01}$	
	No S1 saturation	1.00	
	S2 pulse shape	1.00	
	Minimum S2	$0.99^{+0.01}_{-0.04}$	
	Max S1 fraction per PMT	0.99	
	Prompt LSV	0.95	
	Delayed LSV and WCD	0.94	
	Drift time fiducialization		$(36.9 \pm 0.6) \text{ kg}$
Total	$0.82^{+0.01}_{-0.04}$	$(36.9 \pm 0.6) \text{ kg}$	

**82% global acceptance, dominated by veto cuts**

**20% of exposure sacrificed to drift time fiducialization**

**Mass uncertainty dominated by Teflon contraction at 87K**

# ANALYSIS CUTS: PUTTING IT ALL TOGETHER

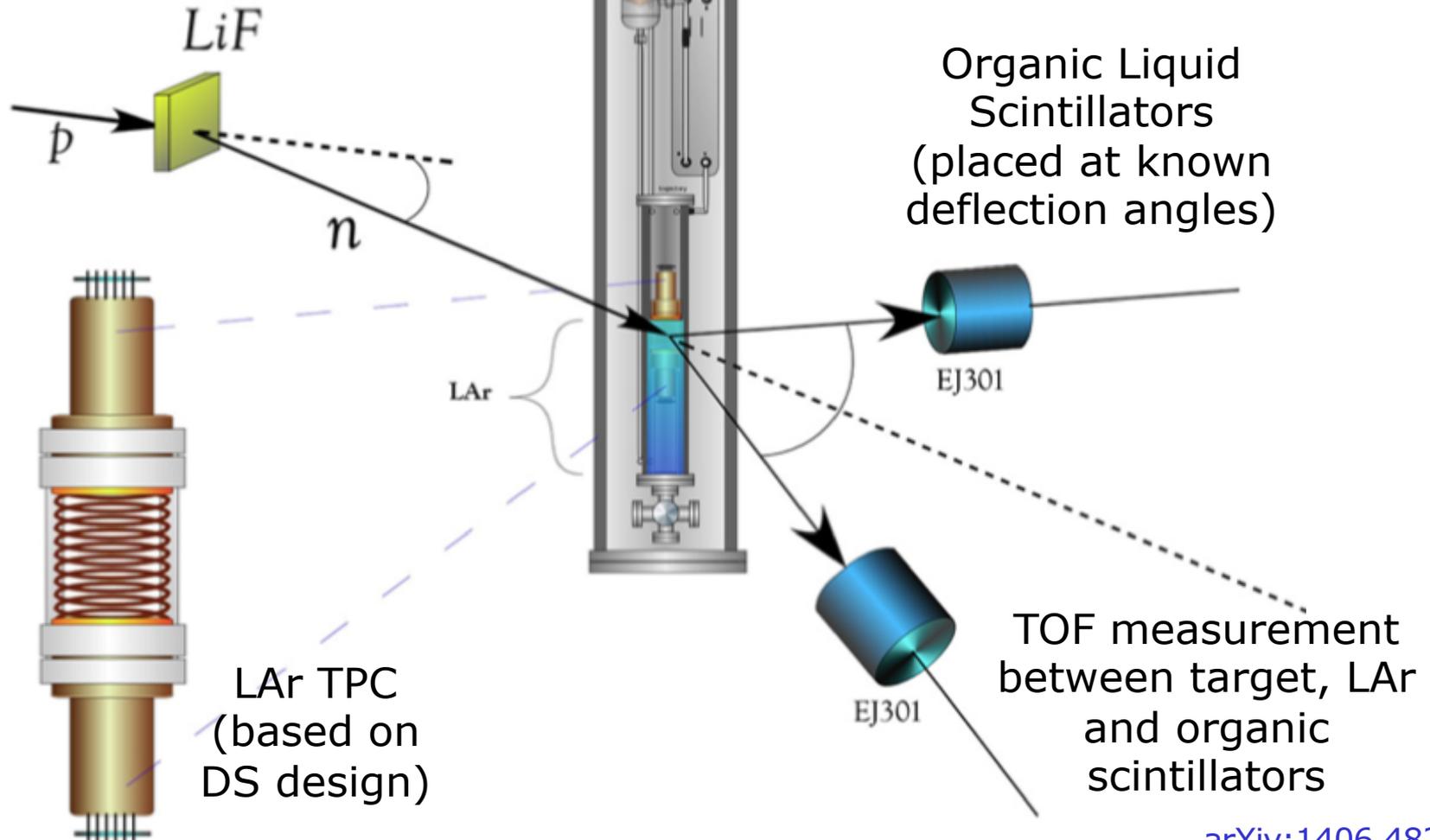
	Cut	Residual Livetime	Acceptance	Fiducial Mass
Run	Usable runs	$(53.8 \pm 0.2) \text{ d}$		
	Automated selection	$(51.1 \pm 0.2) \text{ d}$		
	Single run	$(48.8 \pm 0.2) \text{ d}$		
Quality	Baseline found	$(48.8 \pm 0.2) \text{ d}$		
	Time since previous trigger	$(48.7 \pm 0.2) \text{ d}$		
	Large gap	$(48.1 \pm 0.2) \text{ d}$		
	Veto data present	$(47.1 \pm 0.2) \text{ d}$		
Physics	Number of pulses		$0.95^{+0.00}_{-0.01}$	
	First pulse time		$1.00^{+0.00}_{-0.01}$	
	No S1 saturation		1.00	
	S2 pulse shape		1.00	
	Minimum S2		$0.99^{+0.01}_{-0.04}$	
	Max S1 fraction per PMT		0.99	
	Prompt LSV		0.95	
	Delayed LSV and WCD		0.94	
	Drift time fiducialization			$(36.9 \pm 0.6) \text{ kg}$
	Total	$(47.1 \pm 0.2) \text{ d}$	$0.82^{+0.01}_{-0.04}$	$(36.9 \pm 0.6) \text{ kg}$

**Total Exposure for this run:  $(1422 \pm 67) \text{ kg-d}$**

# NEUTRON CALIBRATIONS WITH SCENE

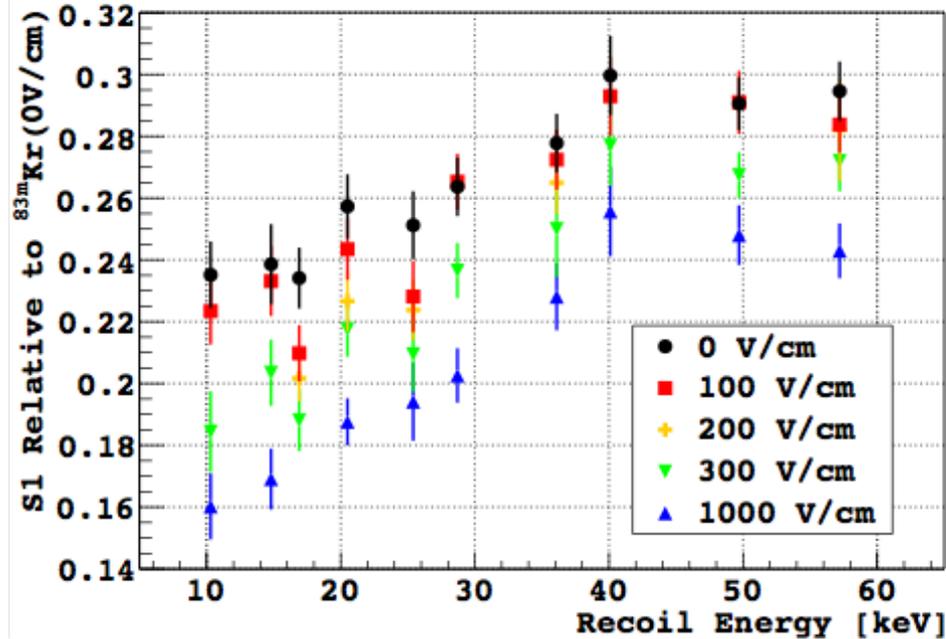
Proton Beam at University of Notre Dame.  ${}^7\text{Li}(p,n){}^7\text{Be}$  reaction produces low-energy monoenergetic neutrons

Scintillation and Ionization  
Efficiency of Noble Elements

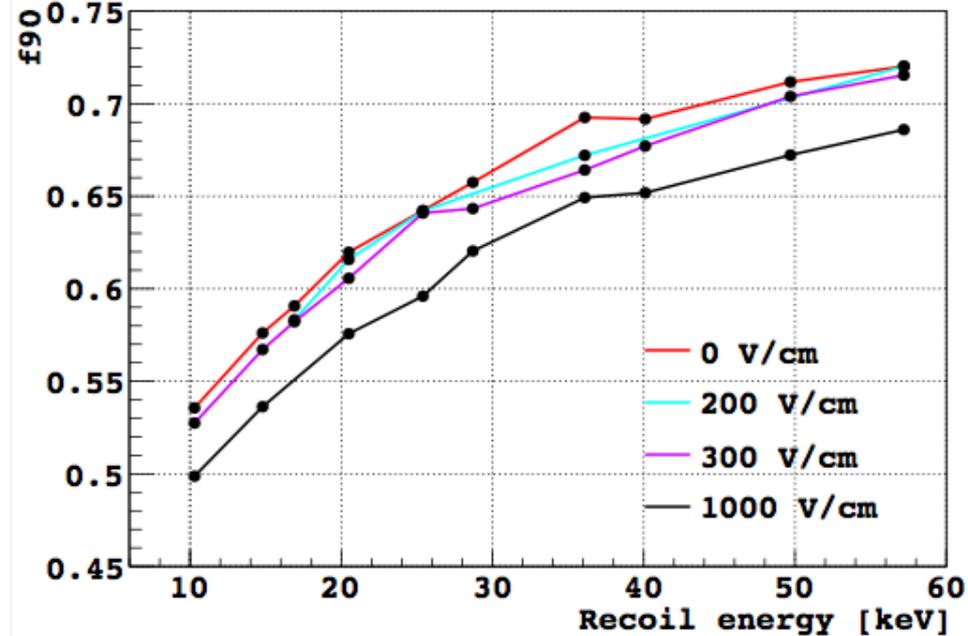


# NEUTRON CALIBRATIONS WITH SCENE

Scintillation yield vs. drift field



Pulse Shape vs. drift field



Source	Energy [keV]	$\mathcal{L}_{\text{eff}, ^{83m}\text{Kr}}$	$S1_{\text{DS-50}}$ [PE]
$^{83m}\text{Kr}$	41.5		$298 \pm 9$
$^7\text{Li}(p,n)$	16.9	$0.202 \pm 0.008$	$27.0 \pm 1.8$
$^7\text{Li}(p,n)$	20.5	$0.227 \pm 0.010$	$36.8 \pm 2.7$
$^7\text{Li}(p,n)$	25.4	$0.224 \pm 0.010$	$45.0 \pm 3.3$
$^7\text{Li}(p,n)$	36.1	$0.265 \pm 0.010$	$75.7 \pm 5.0$
$^7\text{Li}(p,n)$	57.2	$0.282 \pm 0.013$	$127.6 \pm 9.1$

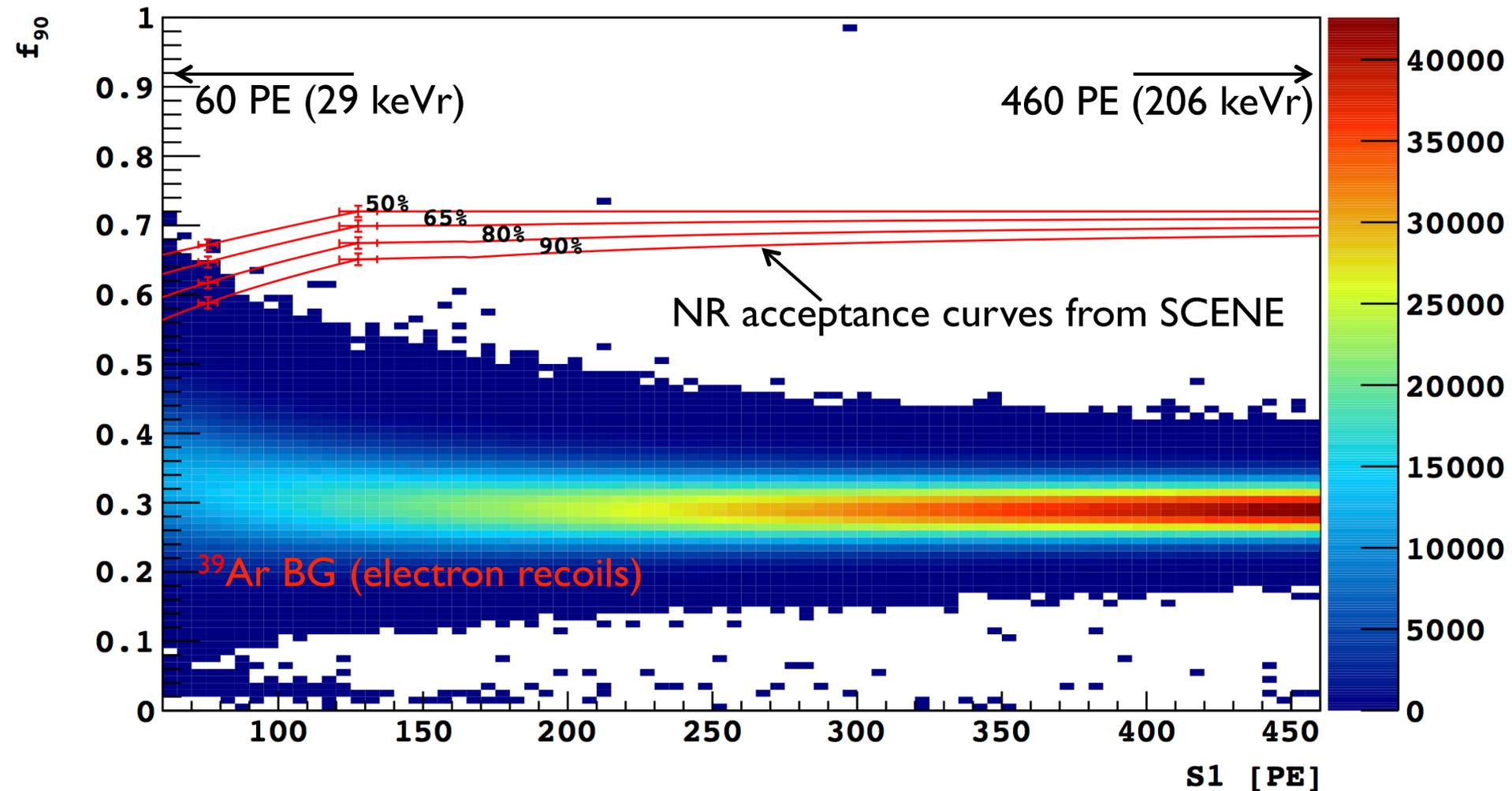
Scintillation yield extrapolated to DS-50 using  $^{83m}\text{Kr}$

Nuclear Recoil  
Acceptance Curves  
also extrapolated  
from SCENE data

Neutron source  
calibration in DS-50  
just completed

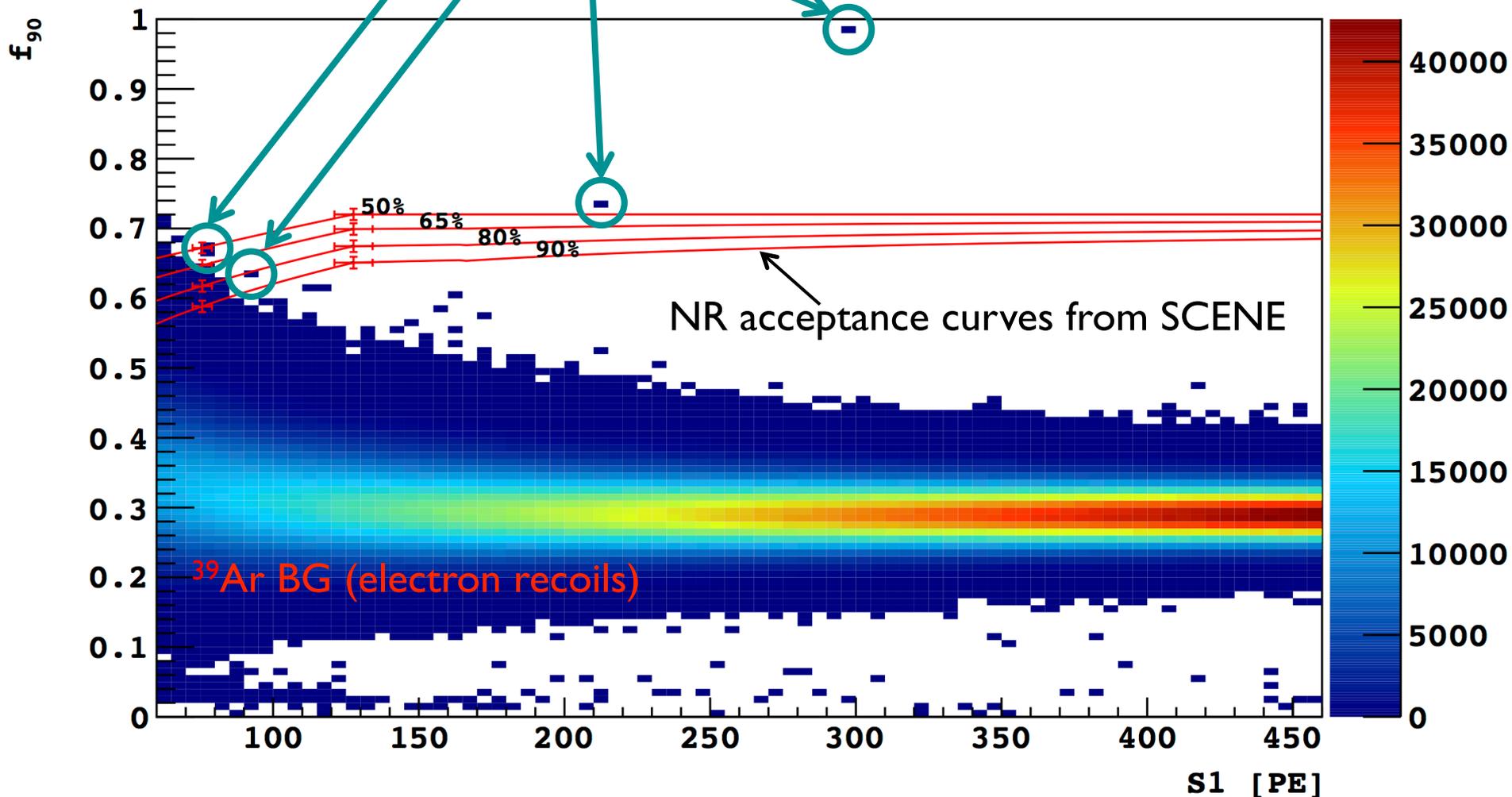
# WIMP SEARCH BOX (1422 ± 67) kg·d

WIMP search plot. All analysis cuts except for veto.



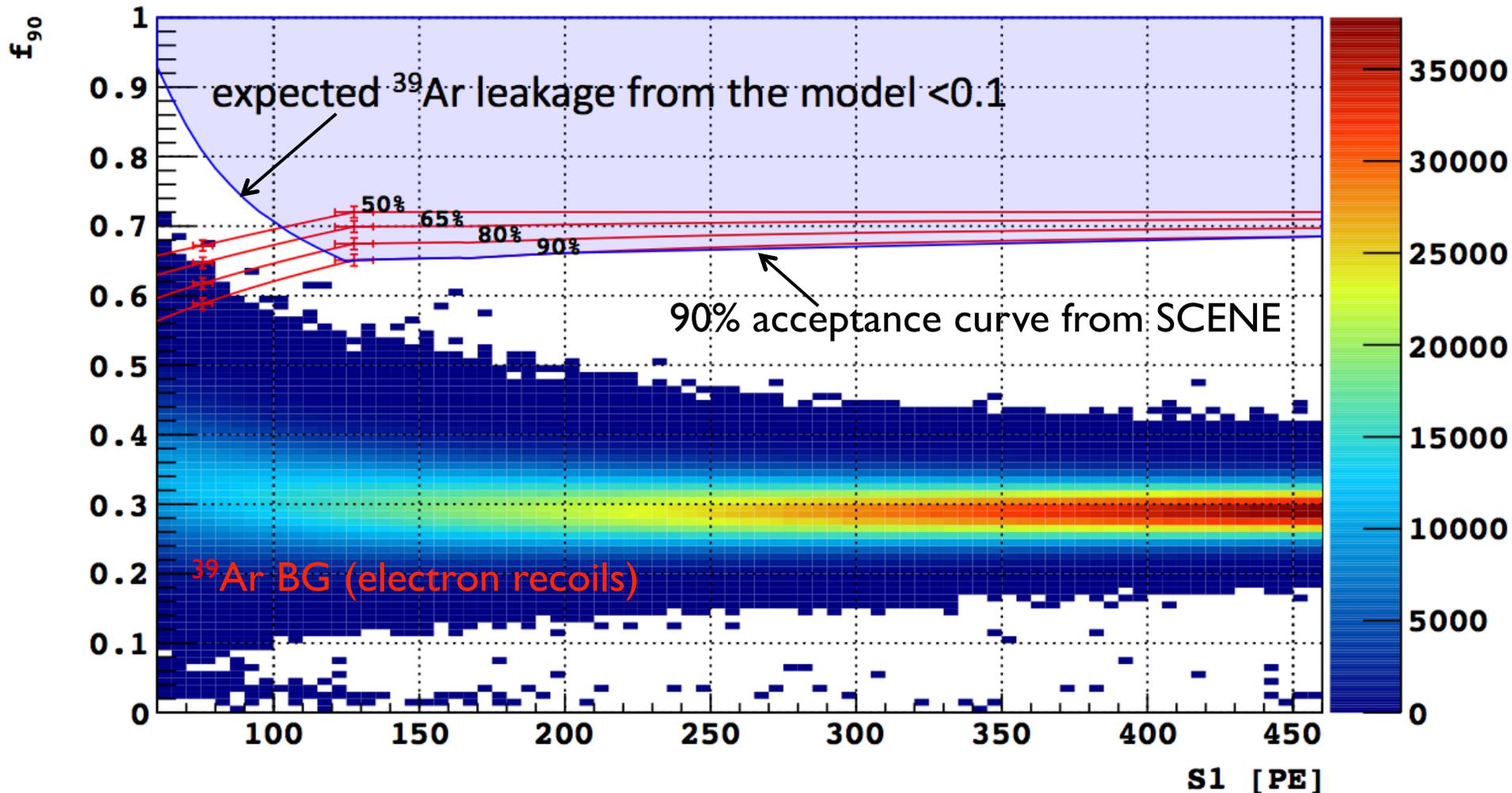
# WIMP SEARCH BOX (1422 ± 67) kg·d

Neutron candidates tagged by Veto.



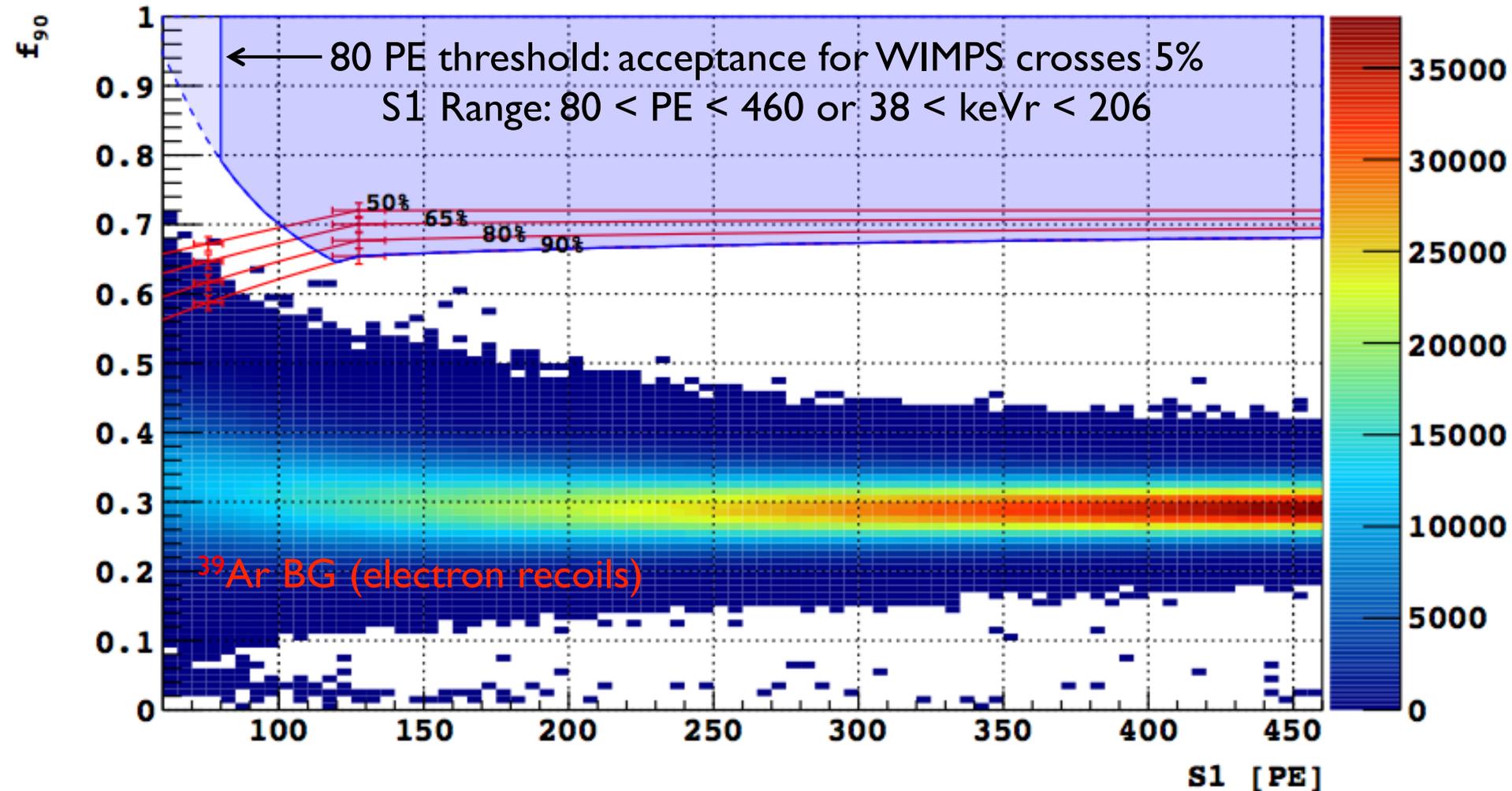
# WIMP SEARCH BOX (1422 ± 67) kg·d

WIMP search box. All analysis cuts applied in this plot.



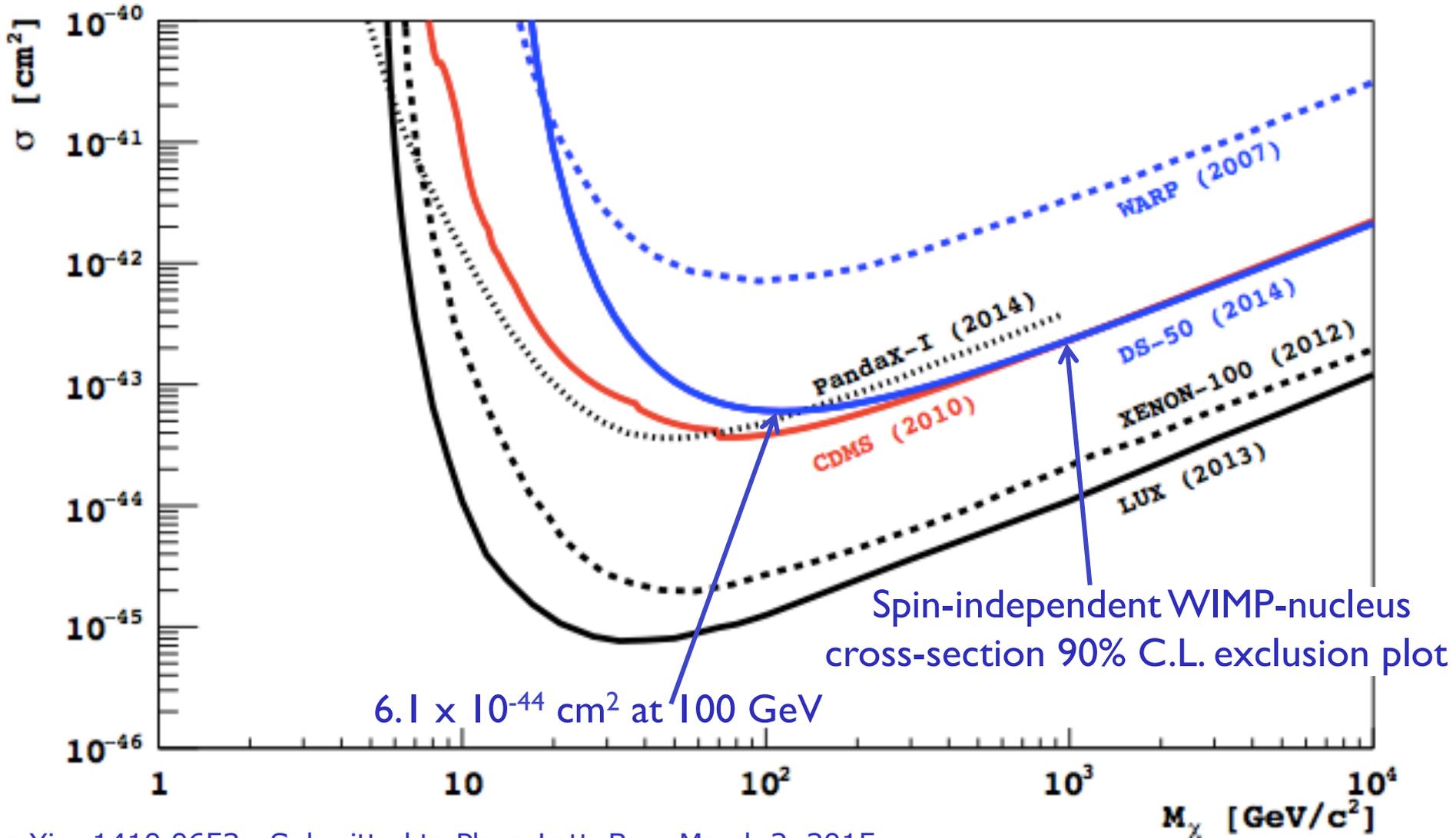
# WIMP SEARCH BOX (1422 $\pm$ 67) kg·d

No candidates in search region after all analysis cuts.



# EXCLUSION PLOT FROM DS-50

Most sensitive DM search ever performed with an argon target



# SUMMARY AND OUTLOOK

---

- **Background-free exposure of  $(1422 \pm 67)$  kg·d**
  - $^{39}\text{Ar}$  background in a  $\sim 20$ -years UAr run would be rejected
  - Future ton-scale LAr TPCs can be free of  $^{39}\text{Ar}$  background
  - Most sensitive DM search ever performed with an argon target
  - Only background-free detector to reach  $10^{-44}$   $\text{cm}^2$  sensitivity
- **We finally have better calibrations!**
  - Deploy calibration sources, perform calibration campaign
  - Calibrate veto and TPC using gamma and neutron sources
  - Calibrations completed in February, results coming soon
- **Vastly improved detector sensitivity in 2015**
  - LSV scintillator replaced with low- $^{14}\text{C}$  (fossil) TMB
  - Fill TPC with underground argon in the next weeks
  - Begin dark matter search with underground argon
  - Expected sensitivity for a 3-years run is  $10^{-45}$   $\text{cm}^2$